

# Liquid Argon Imaging Technology

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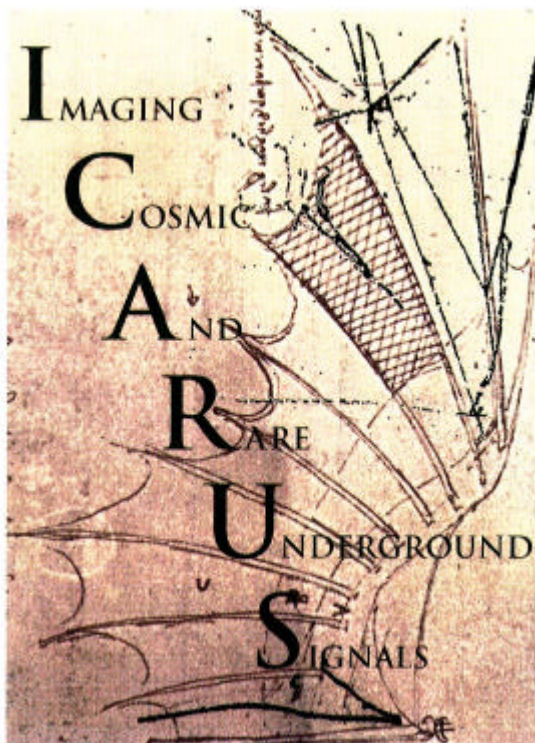
# Images from liquid Argon

- The **liquid Argon Imaging technology**: an **electronic bubble-chamber**, proposed by Carlo Rubbia (CERN-EP/77-08) and supported by INFN over many years of R&D.
- LAr: large sensitive masses, continuously sensitive, self-triggering, 3D views of ionizing events, PID from  $dE/dx$ , range and MS.
- The detector also acts as a performing homogeneous calorimeter of very fine granularity: **ideal device for rare event detection, such as proton decays and neutrino interactions.**
- Argon is inert, not flammable and operates without pressure or moving parts: **particularly suitable in difficult environmental conditions (underground experiments)**

## Present situation

- LAr has recently gained a strong interest in the scientific community, after the:
  - successful operation of the T600 ICARUS detector: the technology is mature
  - realization of the physics potential of high granularity imaging/resolution, for:
    - **underground physics**: proton decay, solar, atmospheric and supernova neutrinos, ...)
    - **short-baseline near detectors**: low/medium cross-sections, high-energy precision neutrino physics
    - **long-baseline** neutrino physics (Super-beams and/or NF)
- One can safely state that:
  - ICARUS at LNGS is the most important milestone for the technology and acts as a full-scale test-bed with a total of 3 kton of liquid Argon in a difficult underground environment
  - **extrapolation to very large LAr detector** (100-200 kton) is under investigation





## ICARUS is...

- a background-free, nucleon decay search experiment
- a solar neutrino experiment
- a supernova watch experiment
- an atmospheric neutrino experiment
- a LBL  $\tau$  and  $e$  appearance experiment in the CNGS beam

The ICARUS Collaboration  
(25 institutes, about 150 physicists)

**ITALY:** L'Aquila, LNF, LNGS, Milano, Napoli, Padova, Pavia, Pisa, CNR Torino, Torino Univ., Politec. Milano.

**SWITZERLAND:** CERN, ETH/Zürich.

**CHINA:** Academia Sinica Beijing.

**POLAND:** Univ. of Silesia Katowice, Univ. of Mining and Metallurgy Krakow, Inst. of Nucl. Phys. Krakow, Jagellonian Univ. Krakow, Univ. of Technology Krakow, A.Soltan Inst. for Nucl. Studies Warszawa, Warsaw Univ., Wroclaw Univ.

**USA:** UCLA Los Angeles.

**SPAIN:** University of Granada, CIEMAT

**RUSSIA:** INR Moscow

# Liquefied rare gases: basic ideas

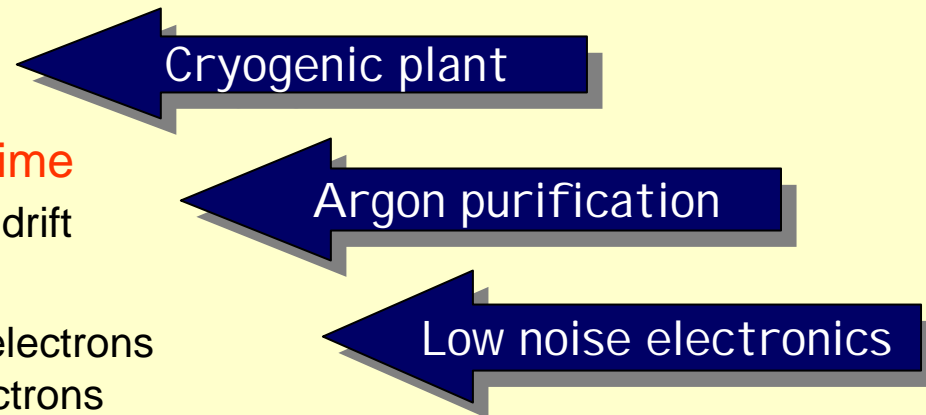
## Ideal materials for detection of ionizing tracks:

- dense, homogeneous, active target
- do not attach electrons (long drift paths possible in liquid phase)
- high electron mobility (quasi-free drift electrons, not Neon)
- commercially easy to obtain (in particular, liquid Argon)
- can be made very pure and impurities freeze out at low temperature
- inert, not flammable
- Ar constitutes 0.9% of air content

Element	Density ( $\rho/\text{cm}^3$ )	Energy loss $dE/dx$ (MeV/cm)	Radiation length $X_0$ (cm)	Collision length $\lambda$ (cm)	Boiling point @ 1 bar (K)	Electron mobility ( $\text{cm}^2/\text{Vs}$ )	Cost
Neon	1.2	1.4	24	80	27.1	high&low	
Argon	1.4	2.1	14	80	87.3	500	€
Krypton	2.4	3.0	4.9	29	120	1200	€€
Xenon	3.0	3.8	2.8	34	165	2200	€€€

# Liquid Argon TPC properties

- High density, heavy ionization medium  
 $\rho = 1.4 \text{ g/cm}^3$ ,  $X_0=14 \text{ cm}$ ,  $\lambda_{\text{int}} = 80 \text{ cm}$
- Very high resolution detector  
3D image  $3 \times 3 \times 0.6 \text{ mm}^3$  (400 ns sampling)
- Continuously sensitive
- Self-triggering or through prompt scintillation light
- Stable and safe  
Inert gas/liquid  
High thermal inertia ( $230 \text{ MJ/m}^3$ )
- Relatively cheap detector  
Liquid argon is cheap, it is only “stored” in the experiment  
TPC: # of channels proportional to surface
- Cryogenic temperature  
 $T = 88 \text{ K}$  at 1 bar
- High purity required for long-drift time  
0.1 ppb of  $\text{O}_2$  equivalent for 3 ms drift
- No signal amplification in liquid  
1 m.i.p. over 3 mm yields 20000 electrons  
equivalent noise charge 1200 electrons



# Processes induced by charged particles in liquid Argon

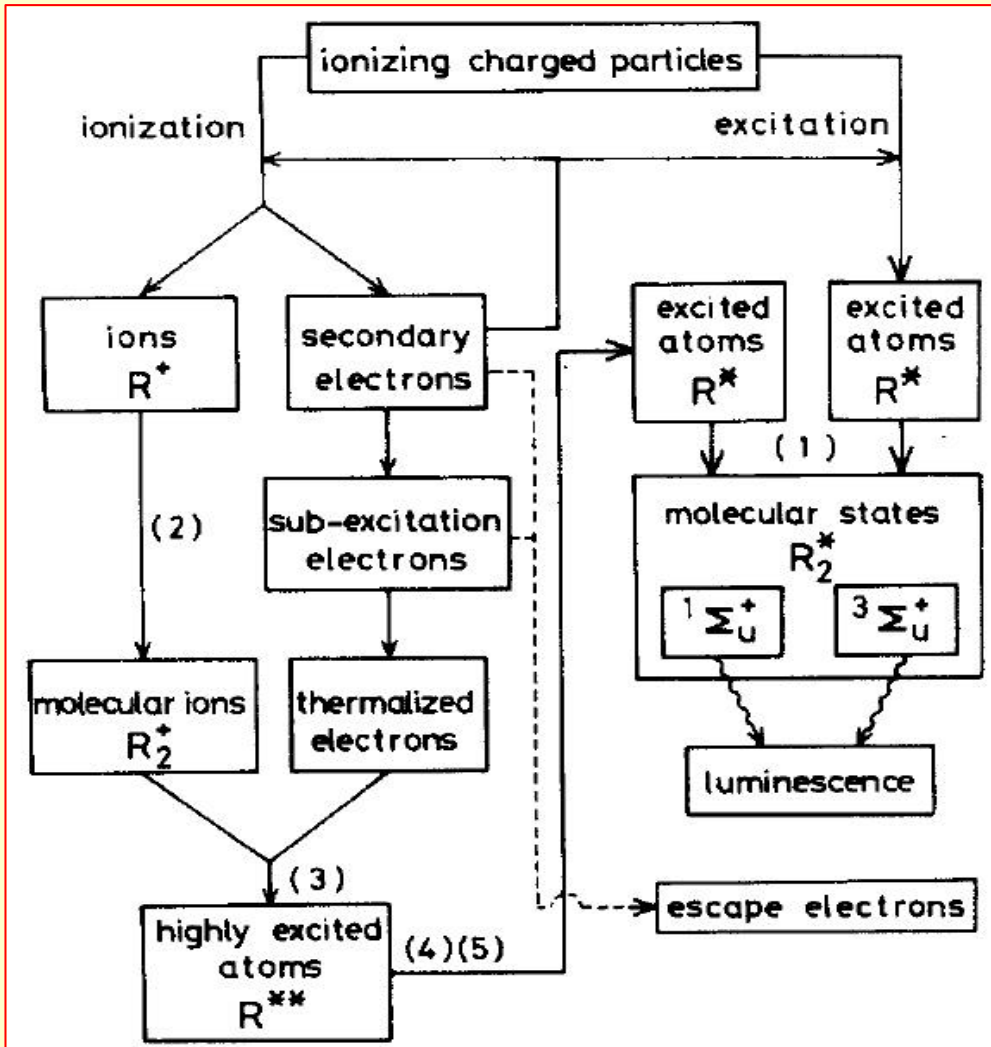
When a charged particle traverses a medium:

- Ionization process
- Scintillation (luminescence)
  - UV spectrum ( $\lambda=128$  nm)
  - Not energetic enough to further ionize (Argon is transparent)
  - Rayleigh-scattering
- Cerenkov light (for fast particle)

→ UV light

→ Charge

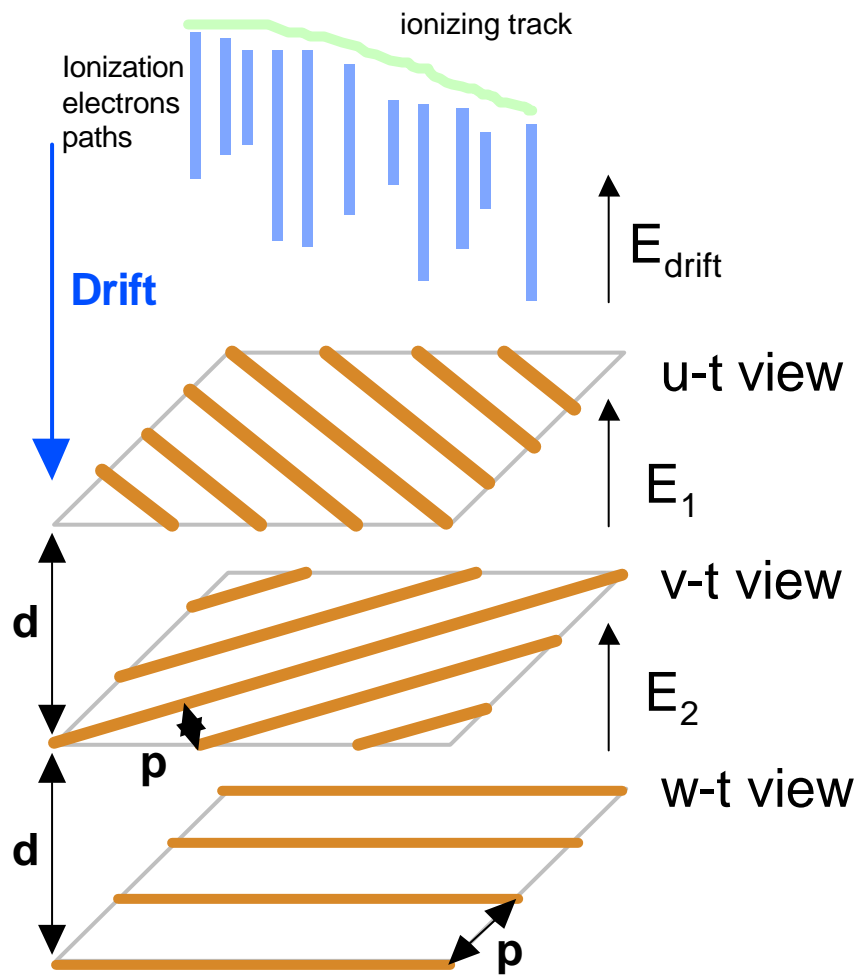
→ Cerenkov light (if  $\beta > 1/n$ )



M. Suzuki et al., NIM 192 (1982) 565

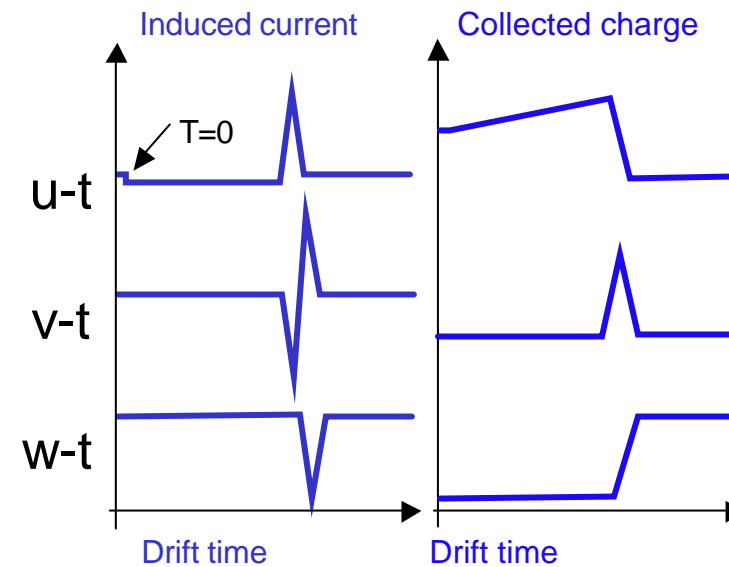
A small addition ( $\sim 500$ ppm) of Xenon shifts the UV radiation to 178 nm

# Charge readout and imaging



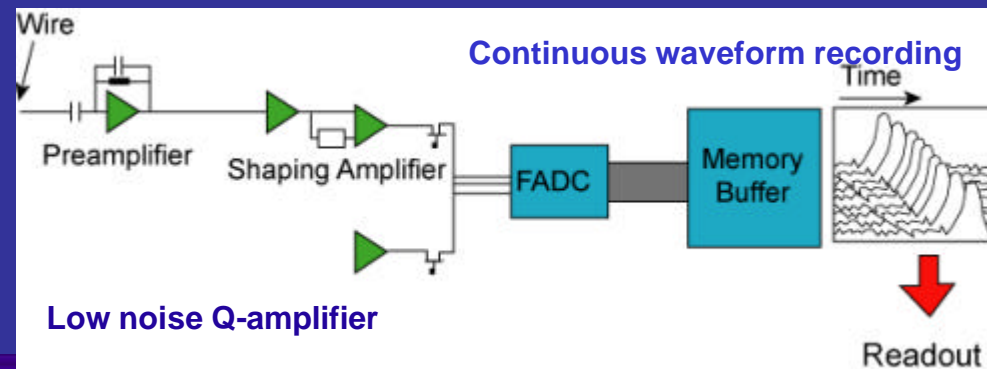
Yield ~ 6000 electrons/mm  
~ 1 fc/mm

In the ICARUS T600 :  $E_{\text{drift}} = 500 \text{ V/cm}$   
 $p = 3 \text{ mm}$   
 $d = 3 \text{ mm}$   
 $r = 0.1 \text{ mm}$

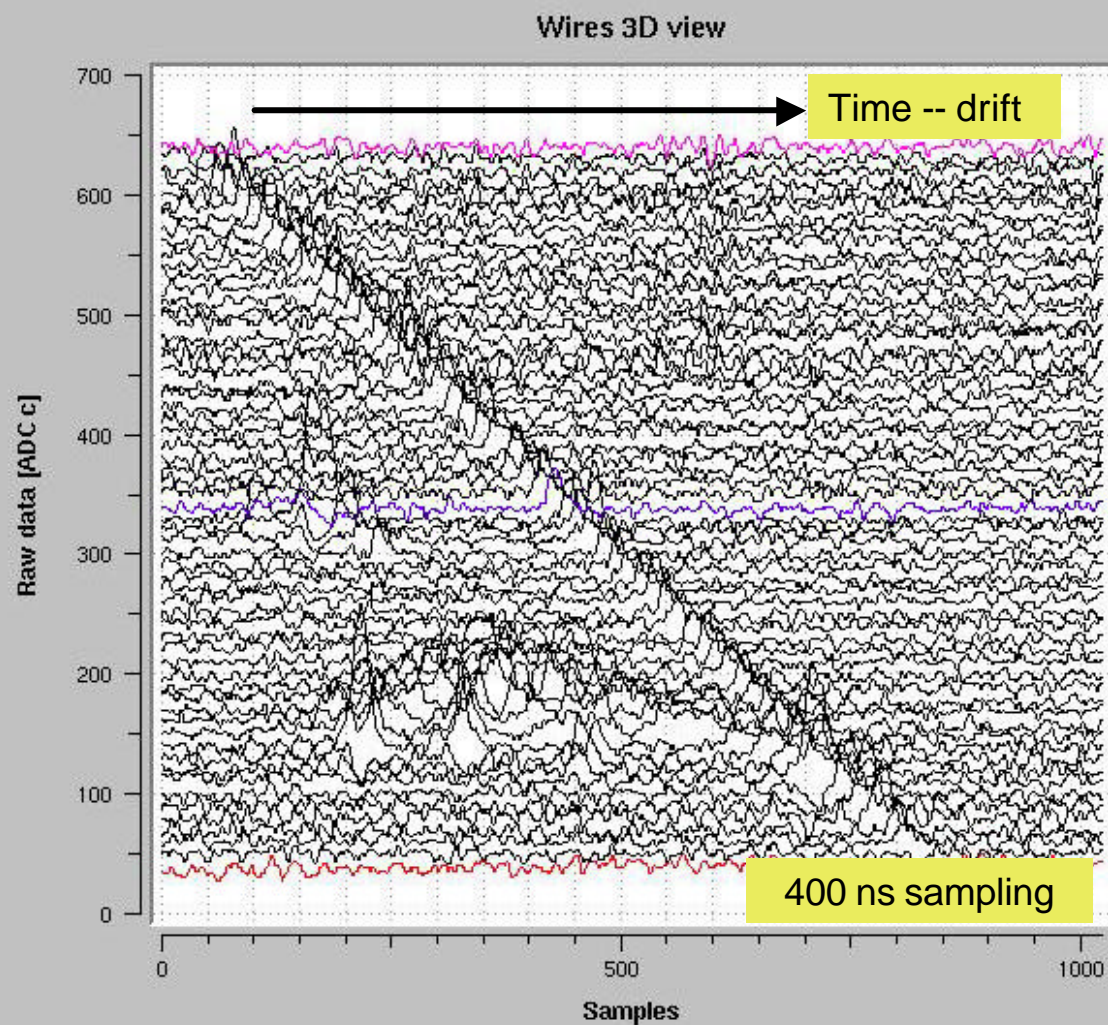




# Principle of signal recording



Wires 3D display

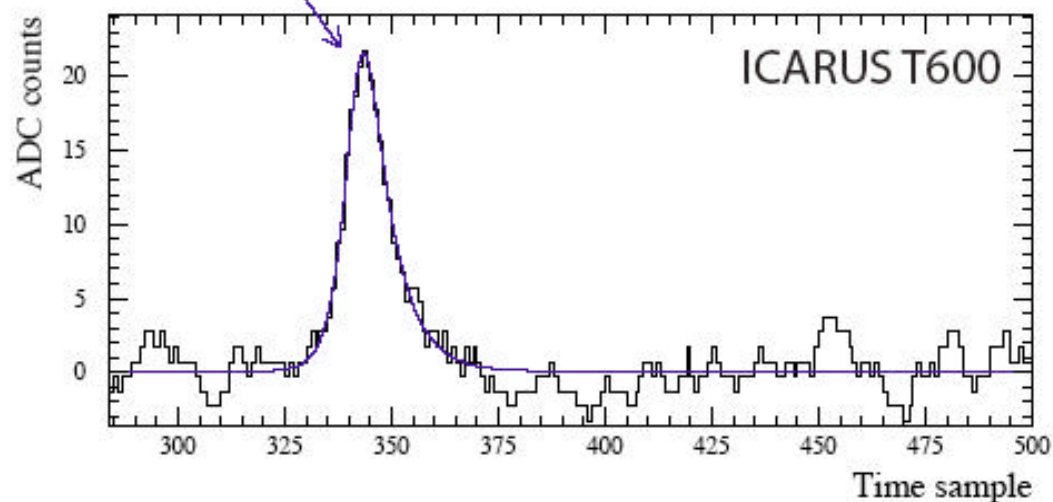
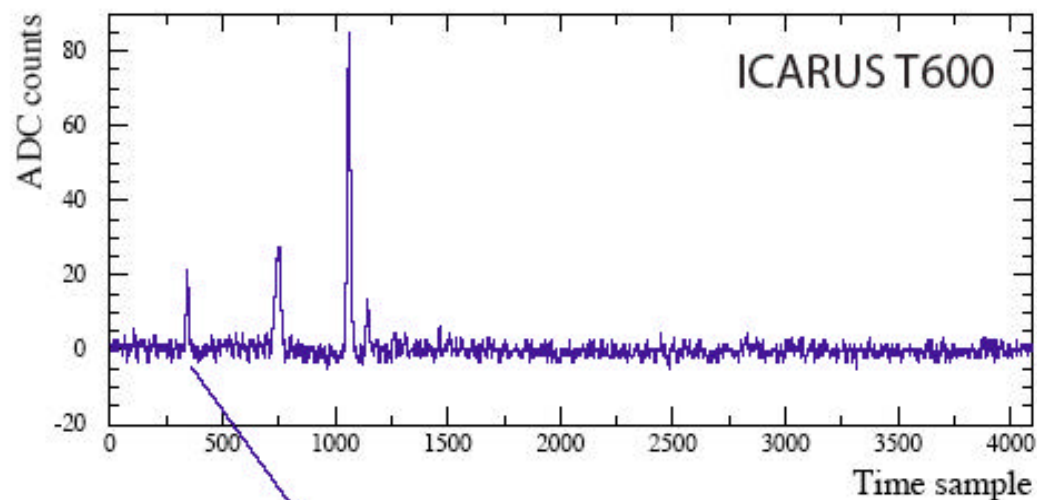


Reconstructed event

40 cm

The plot shows a reconstructed event, which is a single, clear signal trace. A horizontal double-headed arrow at the bottom indicates a width of 40 cm.

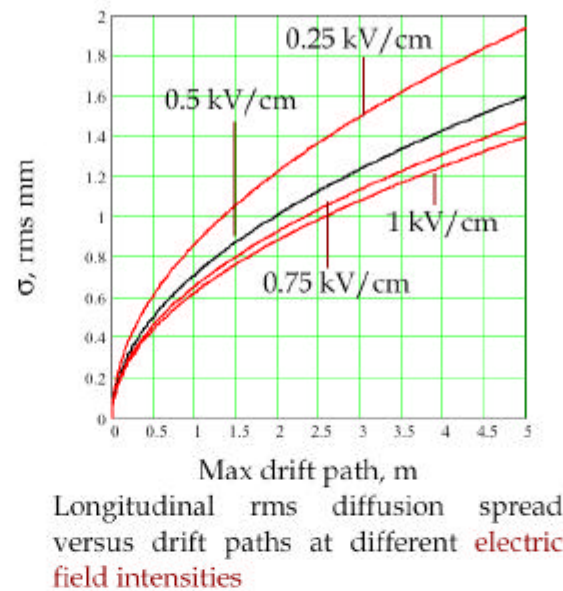
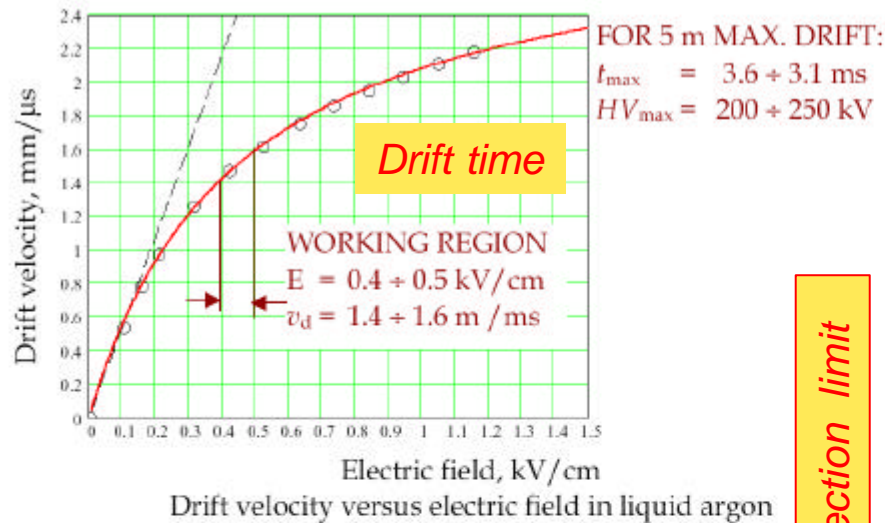
# Single wire signal



Fit function:

$$f(t) = B + A \frac{e^{-\frac{t-t_0}{\tau_1}}}{1 + e^{-\frac{t-t_0}{\tau_2}}}$$

# Electron drift properties in liquid Argon



## R.m.s. diffusion speed

$$\sigma_D = \sqrt{2 \cdot D \cdot \frac{x}{v_d}}$$

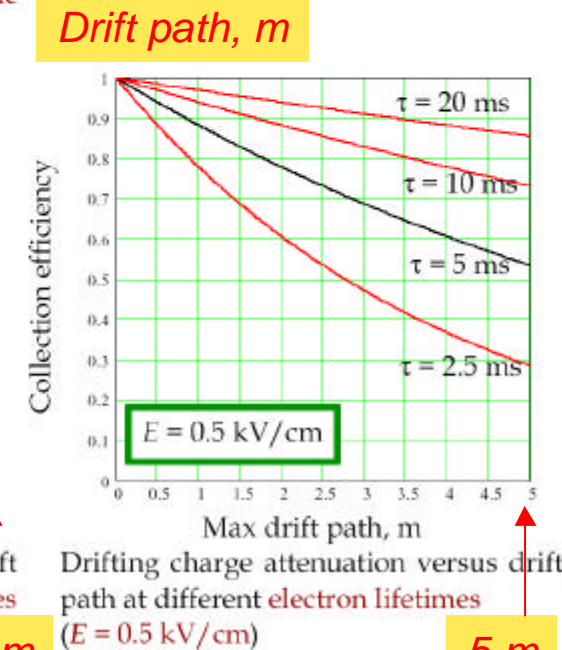
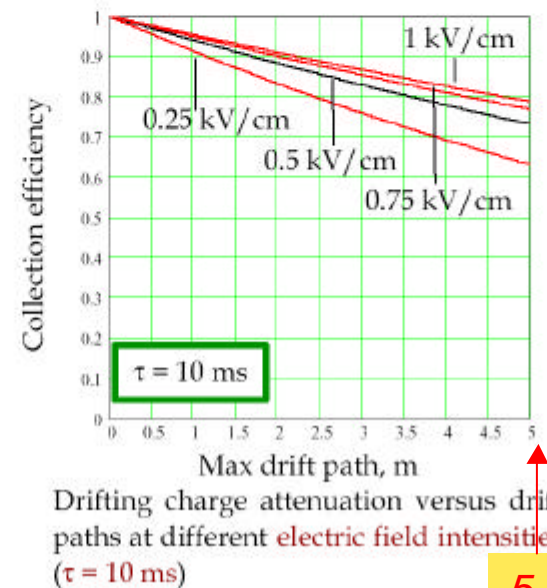
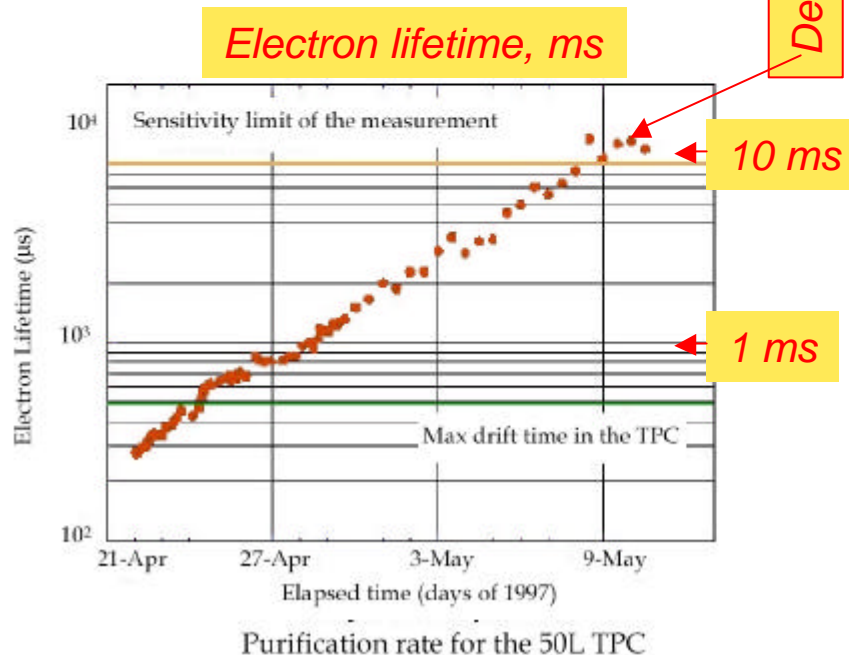
$$D = 4.06 \text{ cm}^2/\text{s}$$

$$\sigma_D = 0.9 \text{ mm} \cdot \sqrt{T_D [\text{ms}]}$$

Longitudinal rms diffusion spread at 0.5 kV/cm

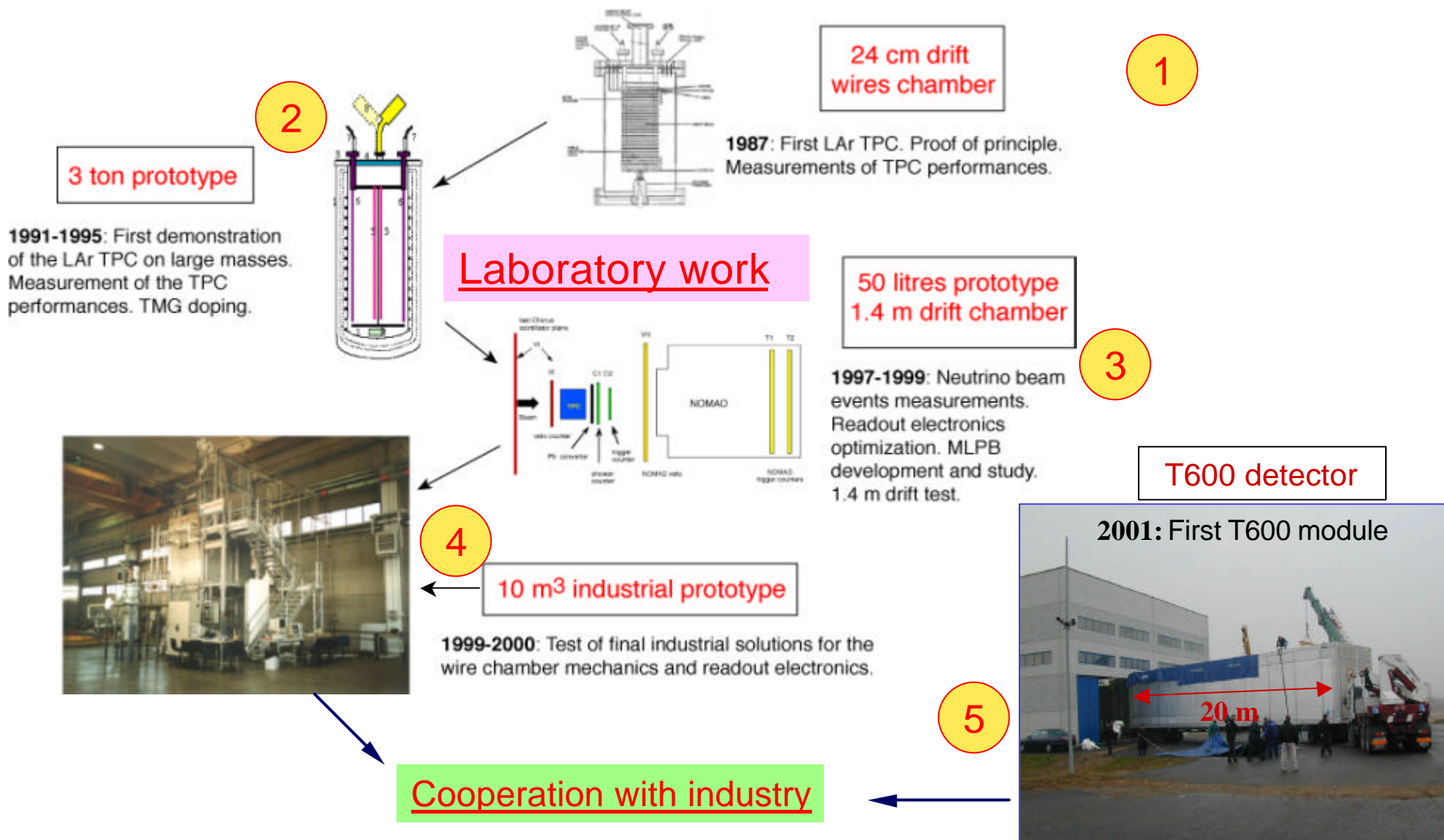
Average  $\langle \sigma_D \rangle = 1.1$  mm

Maximum  $\sigma_{D\max} = 1.6$  mm





# The path to massive liquid Argon detectors



Basic R&D is completed. However, still working on improvements. e.g.:  
TPC prototype for UV-laser calibration and monitoring of LAr purity

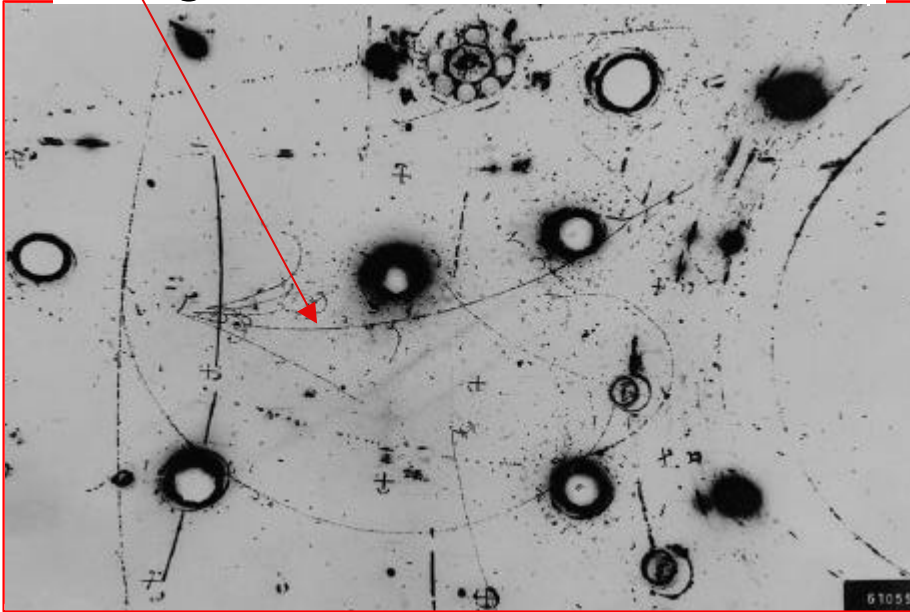




# Thirty years of progress.....

Bubble diameter ~ 3 mm  
(diffraction limited)

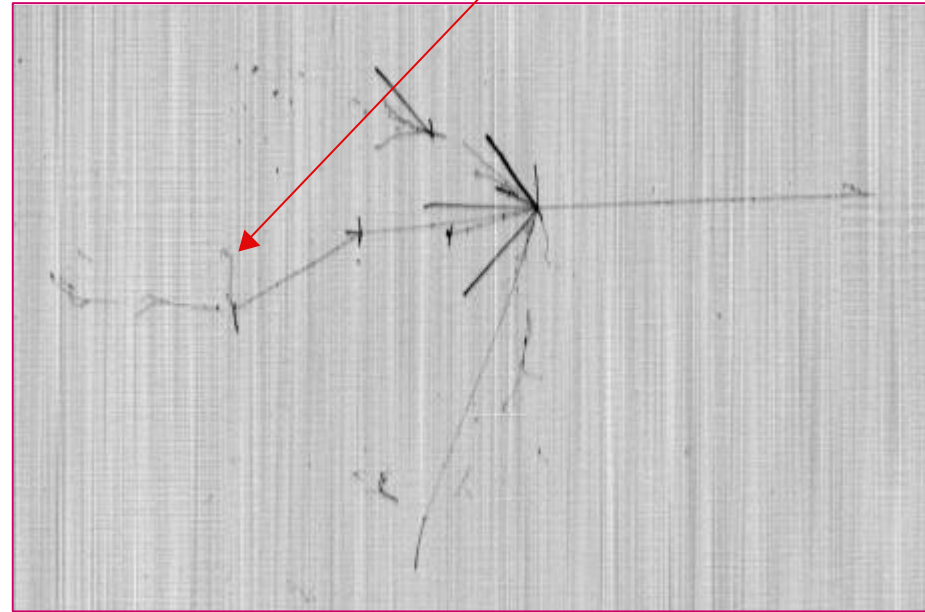
## Gargamelle bubble chamber



Medium	<i>Heavy freon</i>	
Sensitive mass	3.0	ton
Density	1.5	g/cm <sup>3</sup>
Radiation length	11.0	cm
Collision length	49.5	cm
dE/dx	2.3	MeV/cm

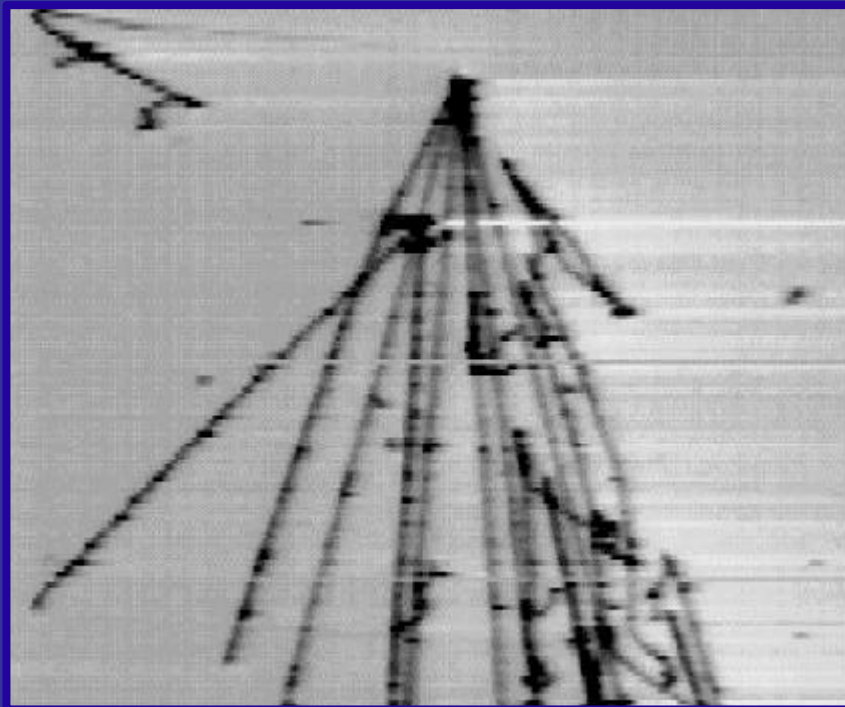
Bubble size ~ 3x3x0.2 mm<sup>3</sup>

## ICARUS electronic chamber

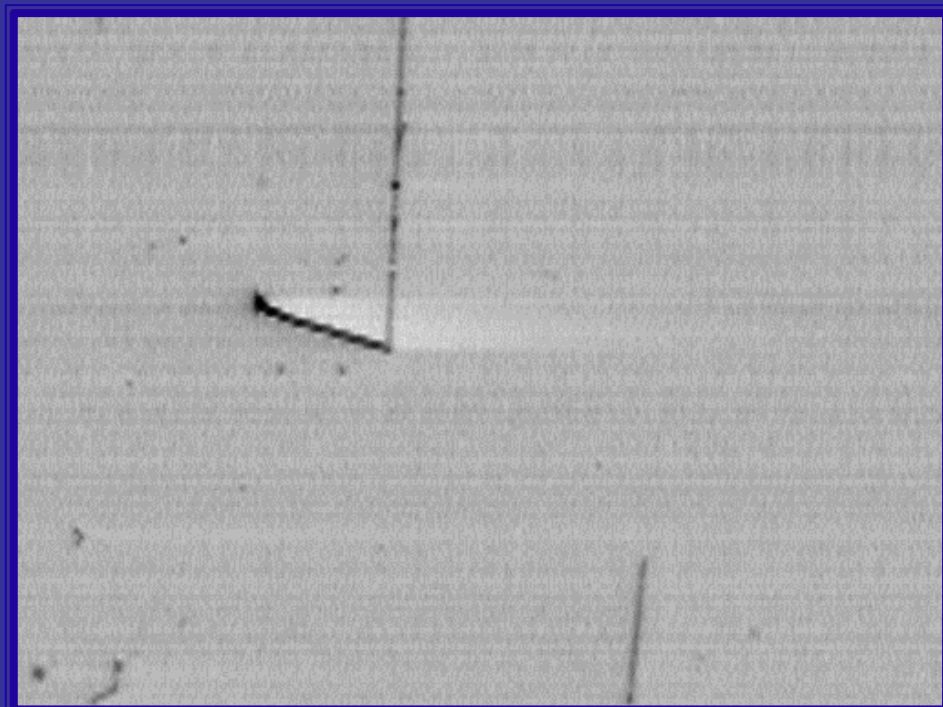


Medium	<i>Liquid Argon</i>	
Sensitive mass	Many	ktons
Density	1.4	g/cm <sup>3</sup>
Radiation length	14.0	cm
Collision length	54.8	cm
dE/dx	2.1	MeV/cm

## 50 l prototype in the CERN WANF neutrino beam

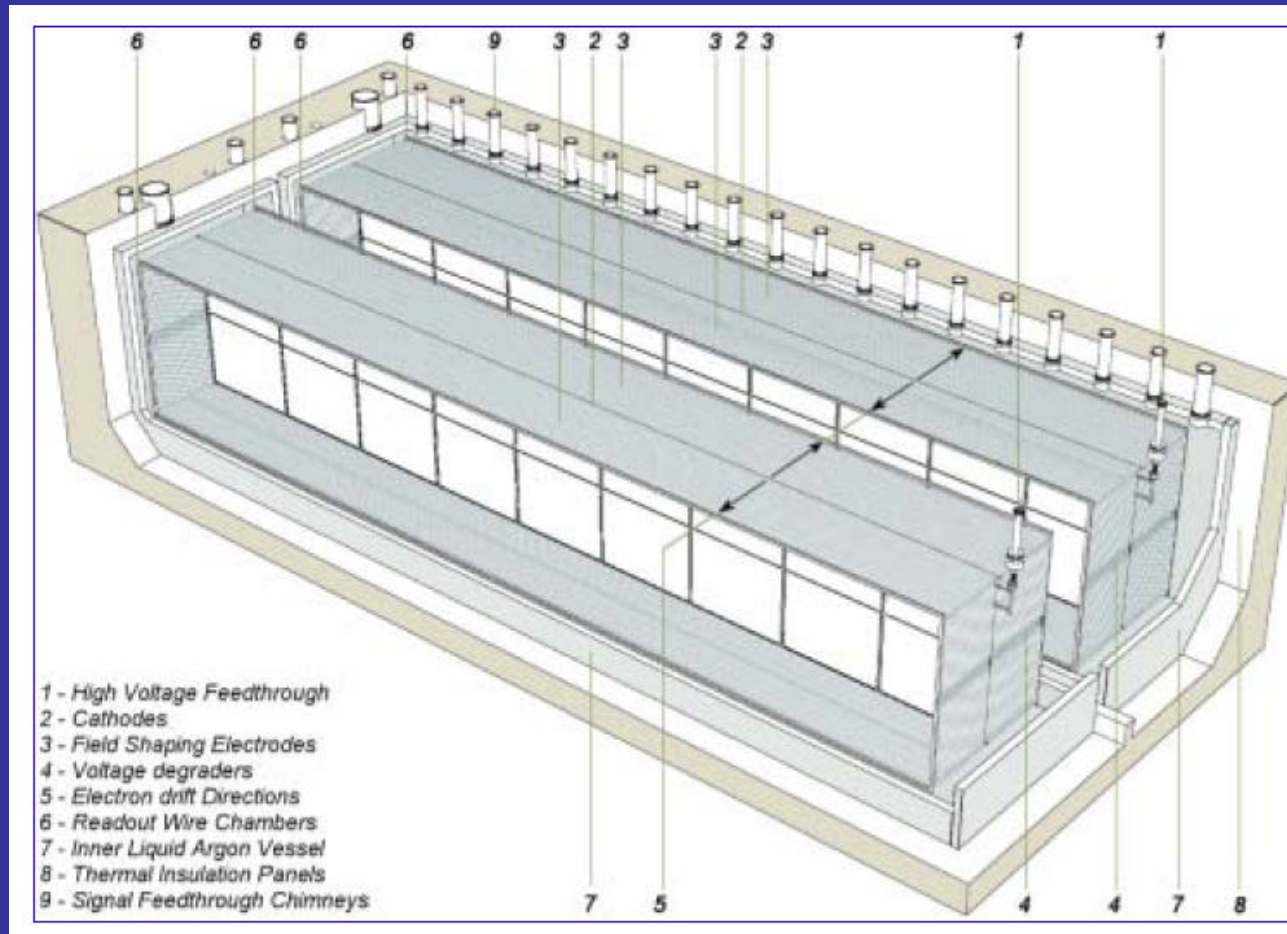


$$\mathbf{n}_m + X \rightarrow \mathbf{m}^- + \text{many prongs}$$



$$\mathbf{n}_m + n \rightarrow \mathbf{m}^- + p$$

# The T600 module



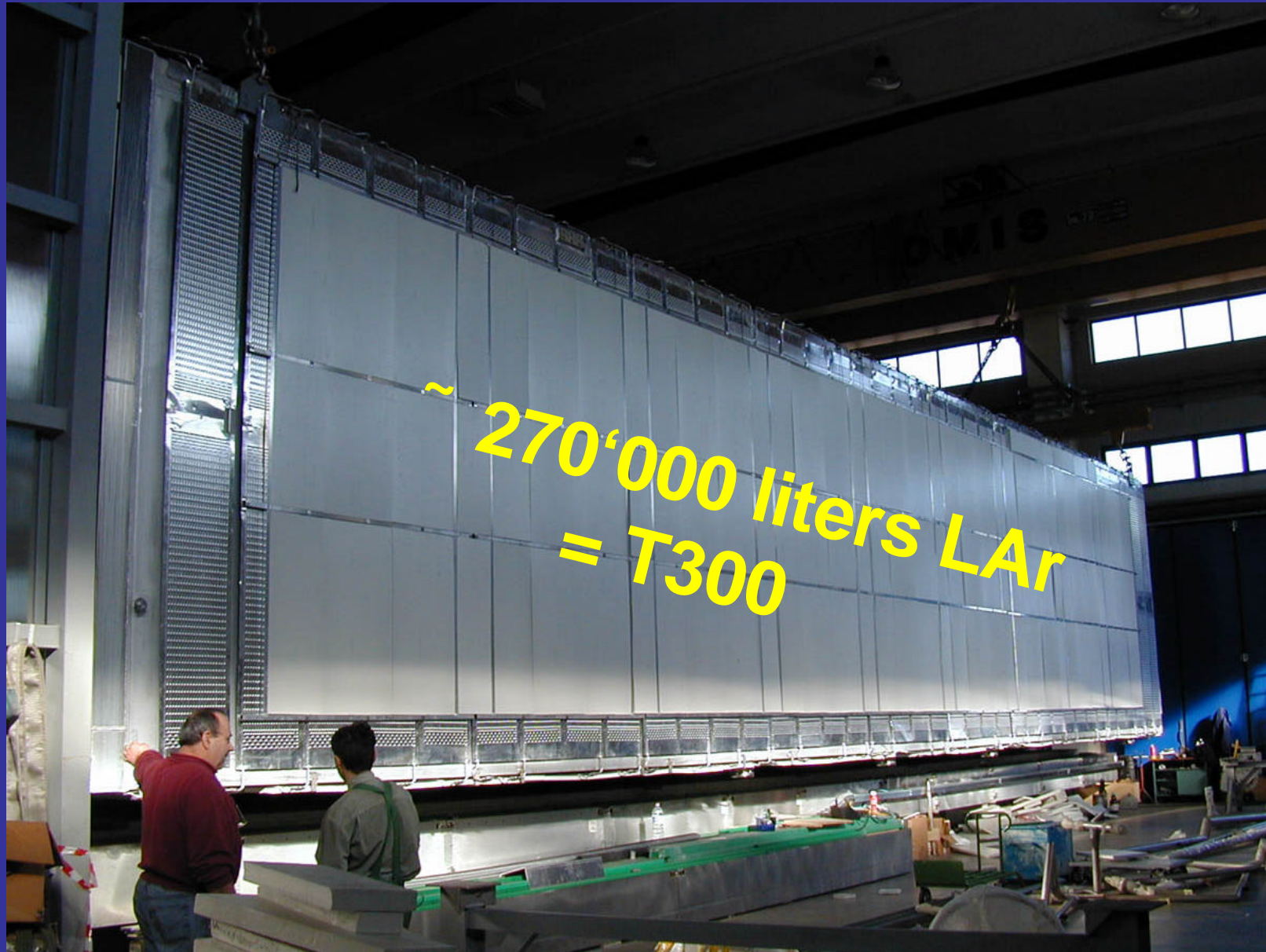
**Design, construction and tests of the  
ICARUS T600 detector.**

To appear on N.I.M.

- Two separate containers
- inner volume/cont. =  $3.6 \times 3.9 \times 19.6 \text{ m}^3$
- 4 wire chambers with 3 readout planes at  $0^\circ, \pm 60^\circ$
- (two chambers/container) ~ 54000 wires
- Maximum drift = 1.5 m  
HV = -75 kV @ 0.5 kV/cm
- Scintillation light readout with 8" VUV sensitive PMTs
- The detector is completely assembled and equipped with the inner detectors, since nearly two years !



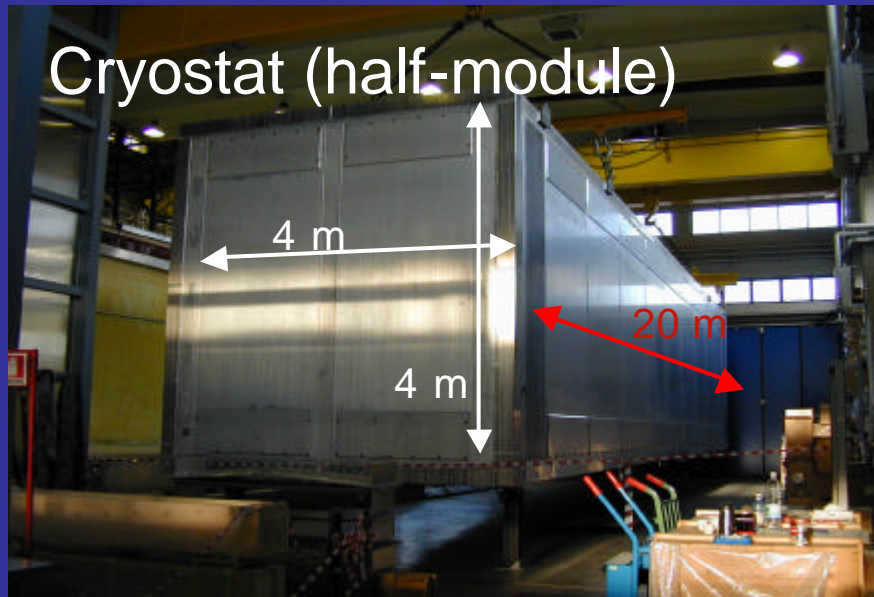
## First T300 cryostat during construction (2001)



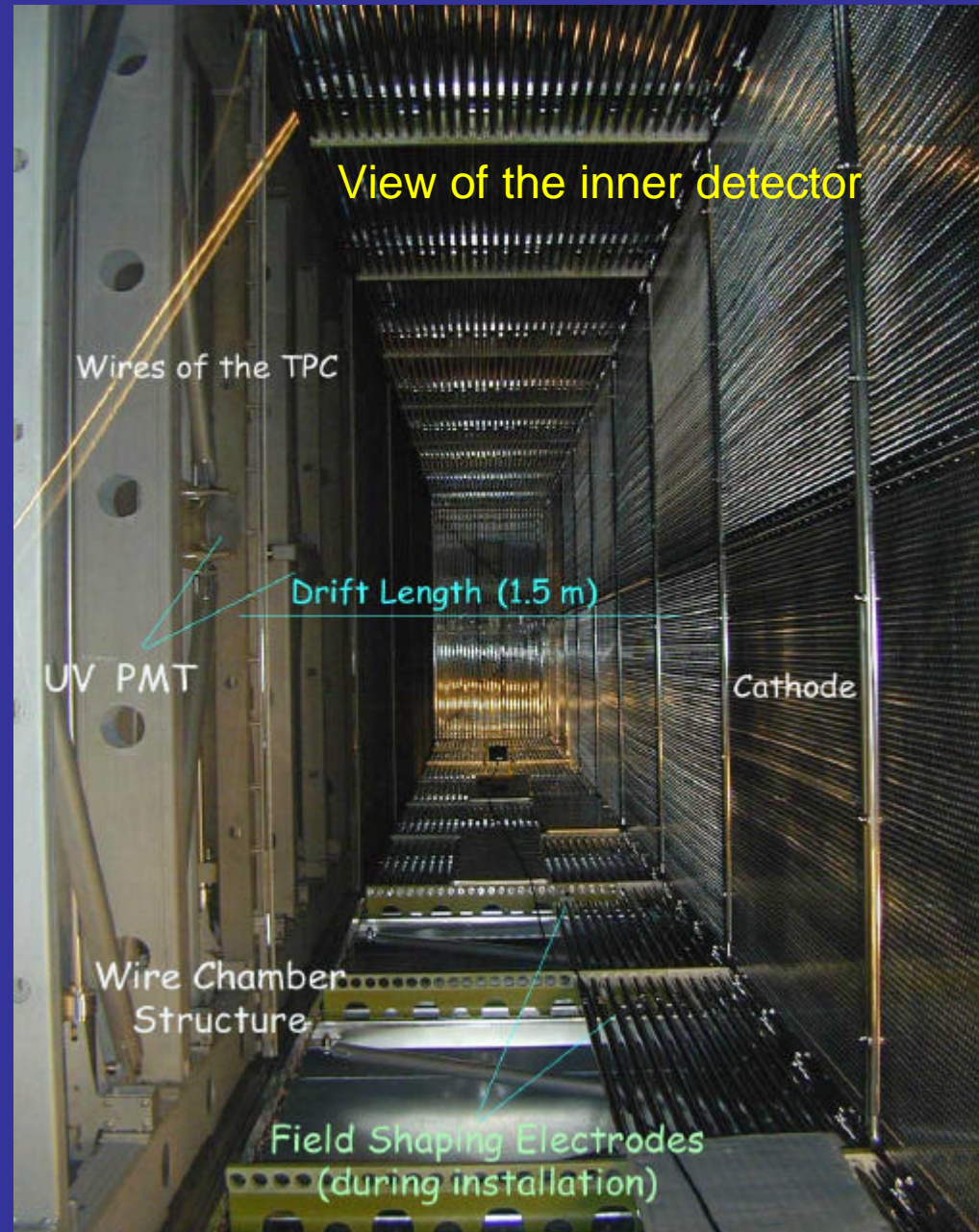
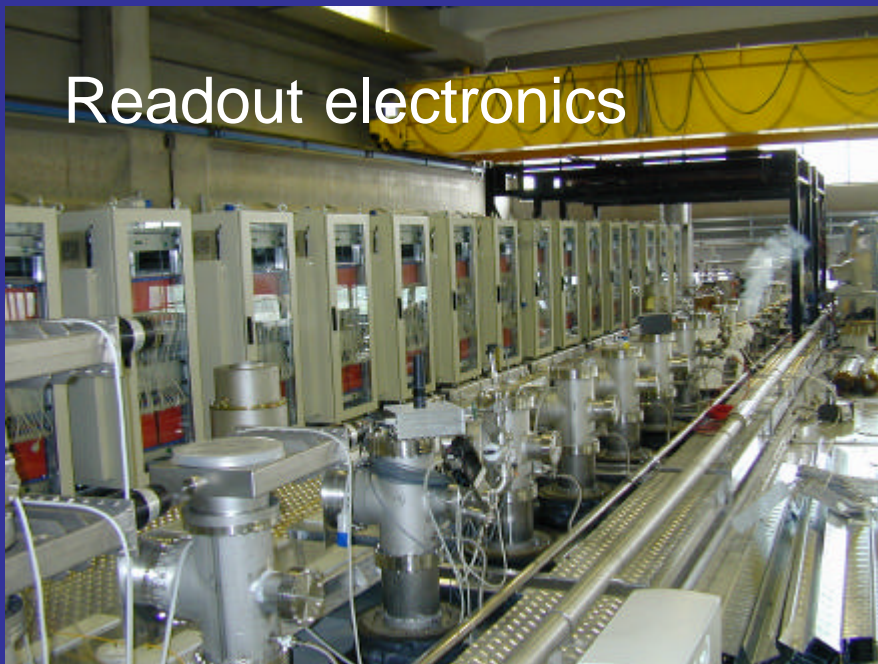


# ICARUS T300 prototype

Cryostat (half-module)



Readout electronics





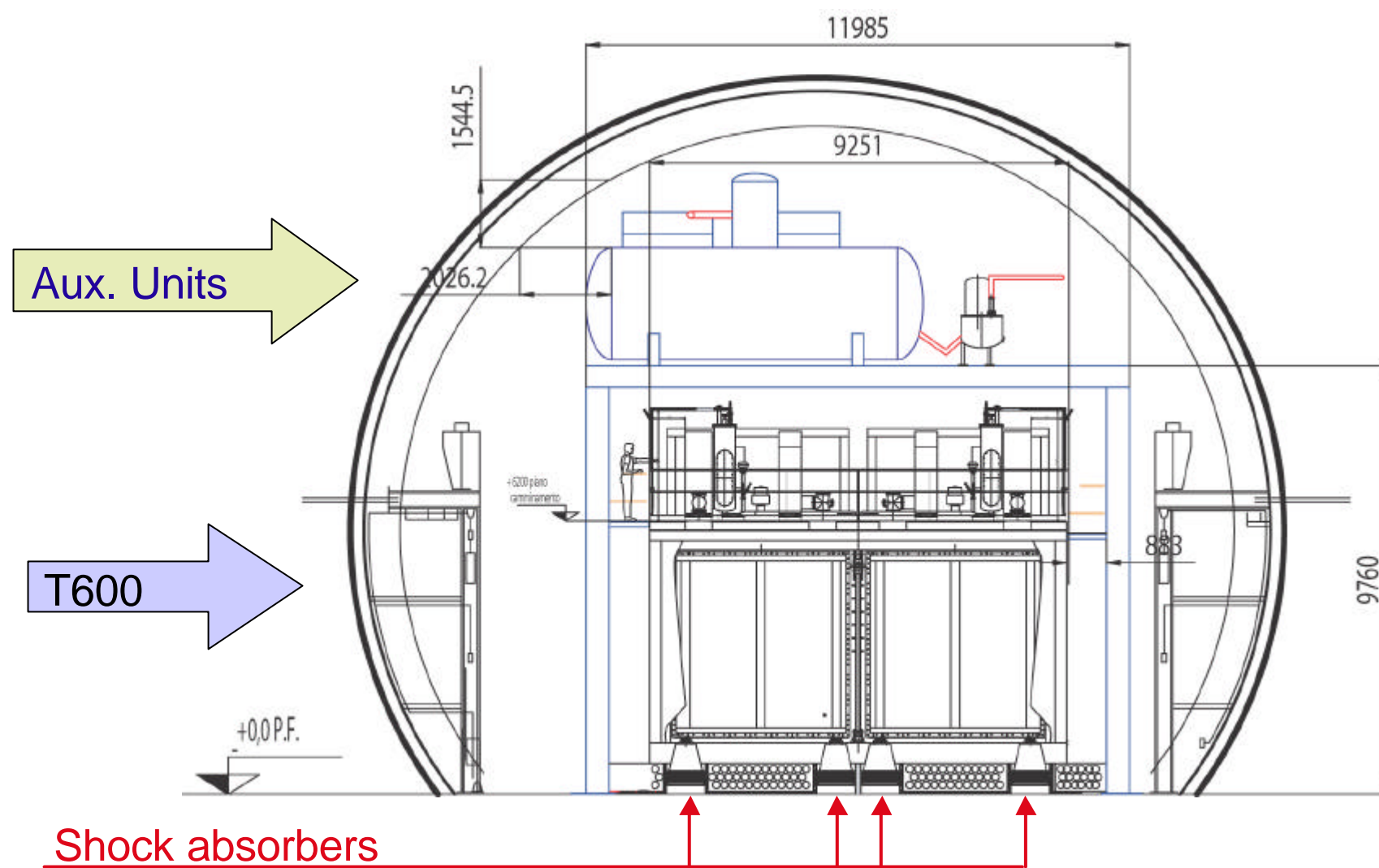




# LNGS Hall B



## First step: T600 installation at LNGS







The T600 installation procedure of the T600 in Hall B has started:

- contract for transportation (Pavia→ LNGS) adjudicated
- interventions on the floor for safety requirements (Government Comm.ner) being defined
- goal: T600 in Hall B by mid 2004 (parking lot position); infrastructure & support structure completed by end 2004; data taking with cosmic events by mid 2005

# Soon, physics with the T600...

In 1 year of T600 running ICARUS will collect about 100 events of this quality (in presence of oscillations)

BG free detection of solar neutrino events ( $E > 8 \text{ MeV}$ )

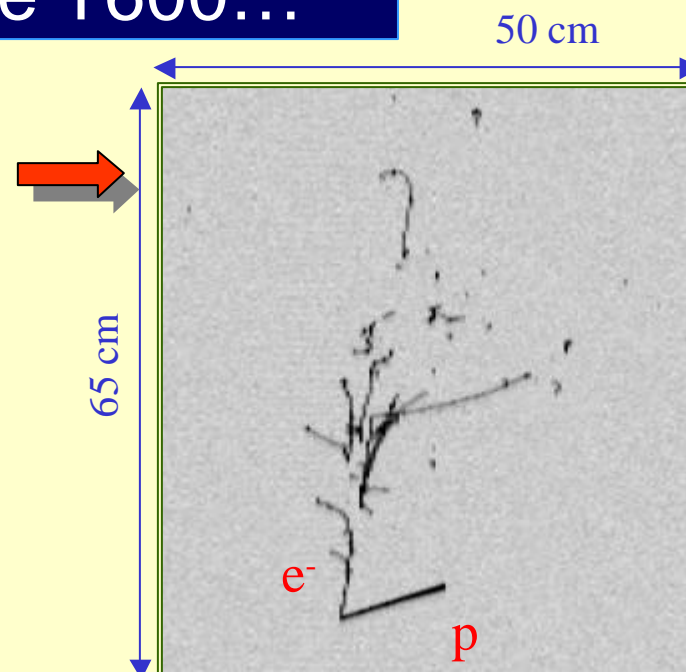
## Solar $\nu$ events per year in T600

Elastic	Fermi	Gamow-Teller
38	165	295

Search for proton decay event topologies

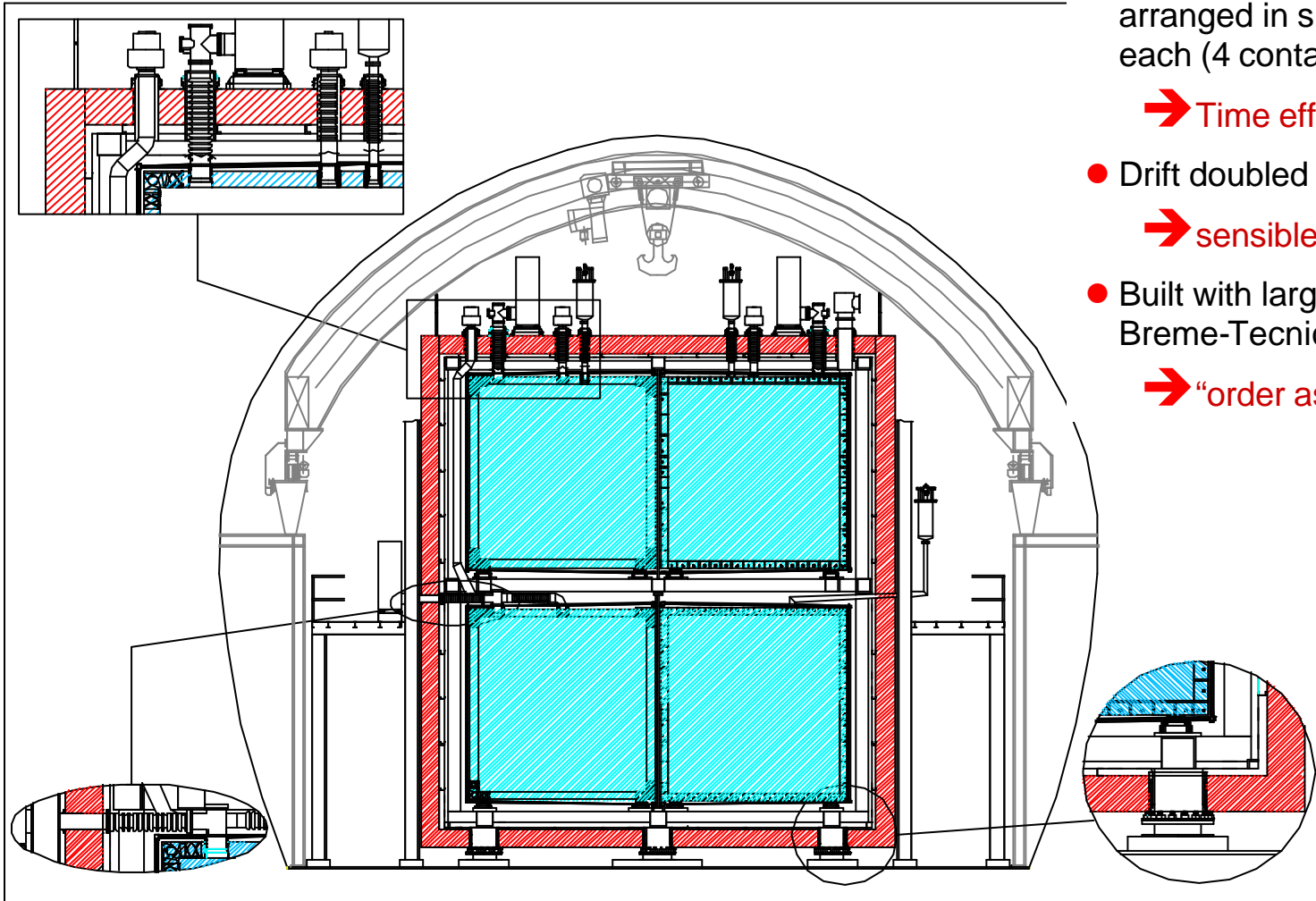
Supernova observatory

..and more: learn about a new technology for astroparticle physics



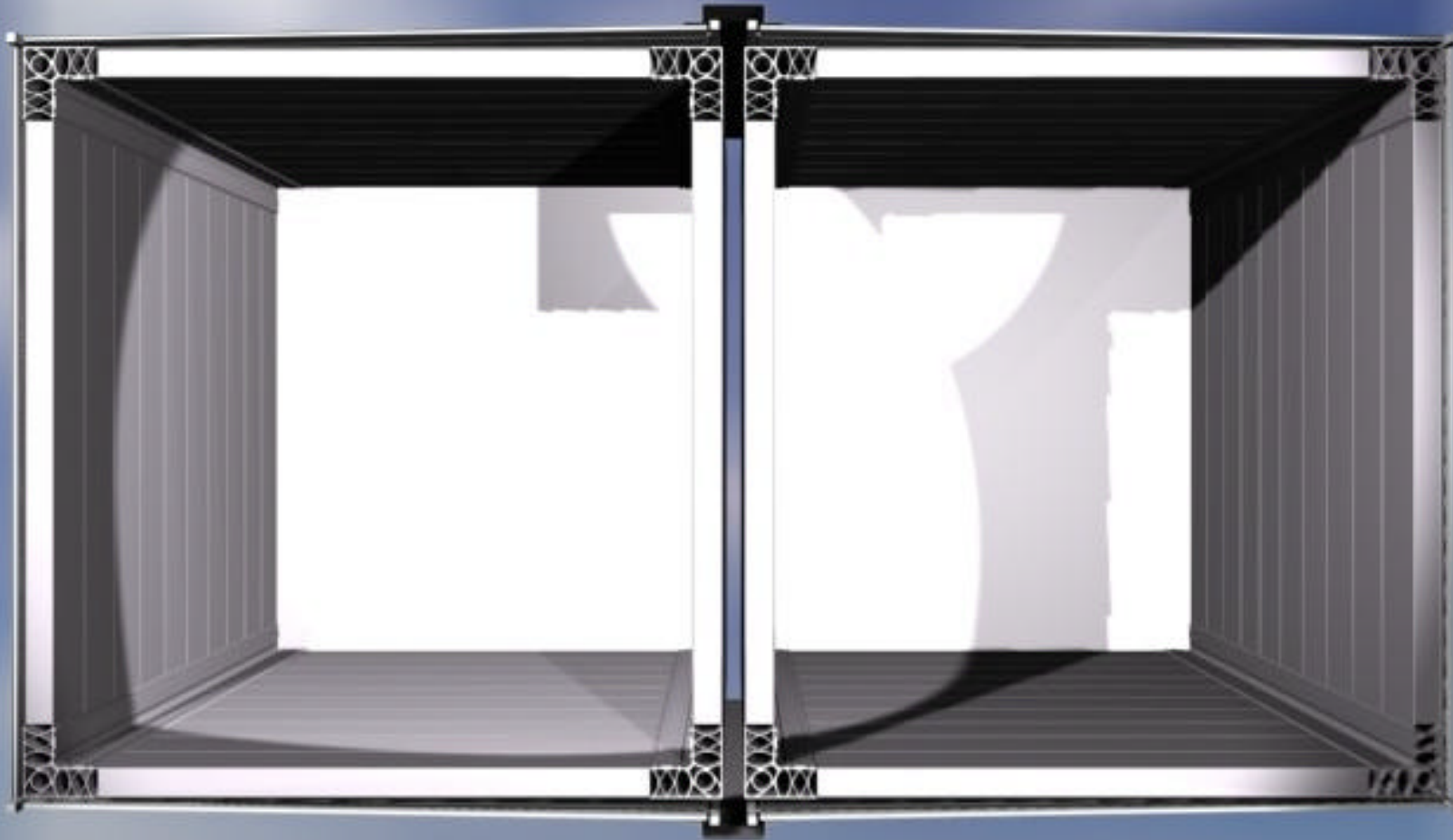


# The ICARUS T1200 “Unit”



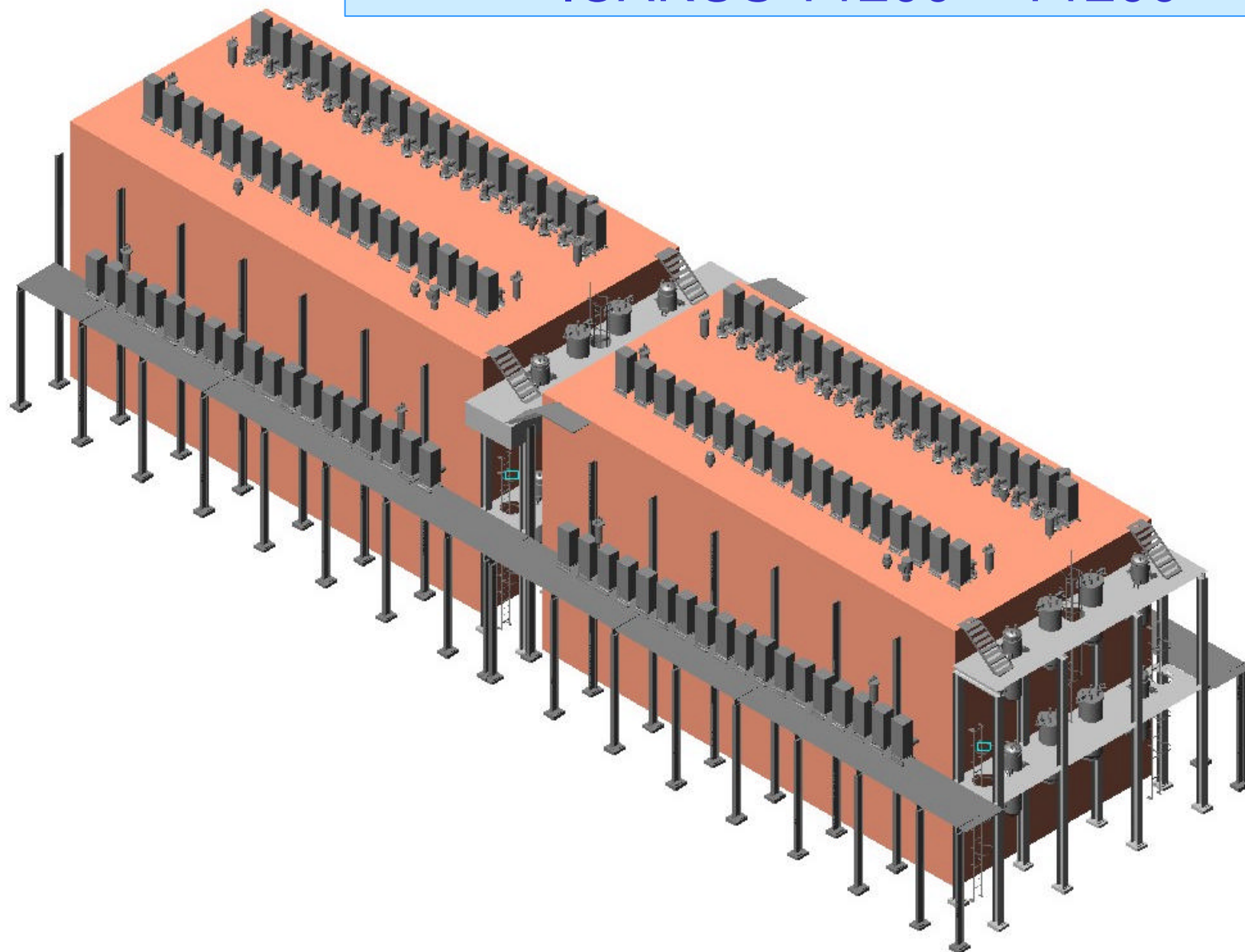
- Based on cloning the present T300 dewars
  - ➔ A cost-effective solution given tunnel access conditions
- Preassembled modules outside the tunnel are arranged in super-modules of about 1200 ton each (4 containers)
  - ➔ Time effective solution (parallelizable)
- Drift doubled 1.5 m ➔ 3 m
  - ➔ sensible solution given past experience
- Built with large industrial support (AirLiquide, Breme-Tecnica, Galli-Morelli, CAEN, ...)
  - ➔ “order as many as you need” solution

Liquid Nitrogen output



Liquid Nitrogen input

# ICARUS T1200 + T1200



## T600 detector performance (surface test run)

- Technical run at surface in 2001: assess the maturity of large scale LAr imaging TPC. Main phases:
  - clean-up (vacuum) 10 days, cool-down 15 days, LAr filling 15 days, debug and data-taking 68 days.
- 18 m long track (request by Scientific Committees) plus a large number of cosmic-ray events:
  - about 28000 triggers with different topologies
  - 4.5 TB of data, 200 MB/event.
- Valuable data: check performance of a such large scale detector. Found that: results of the same quantitative quality as those obtained with small prototypes (e.g. 3 ton, 50 liter, ...) are achieved with a 300 ton device.
- Several papers published with test data →

**Analysis of the liquid Argon purity in the ICARUS T600 TPC.**

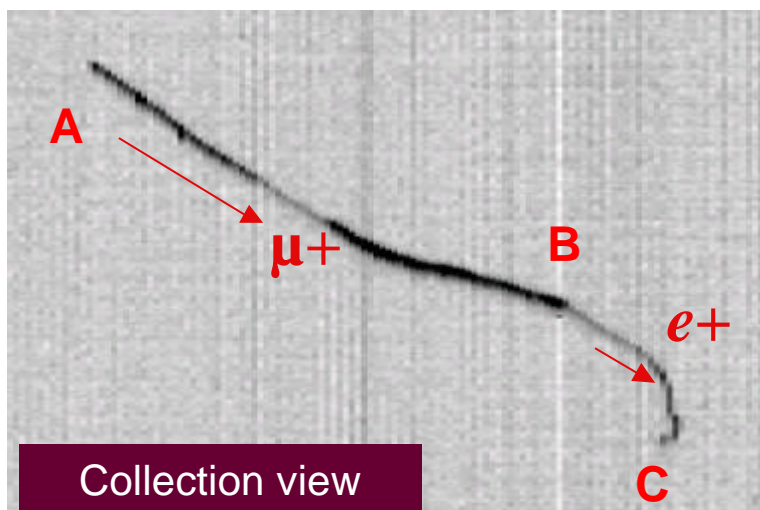
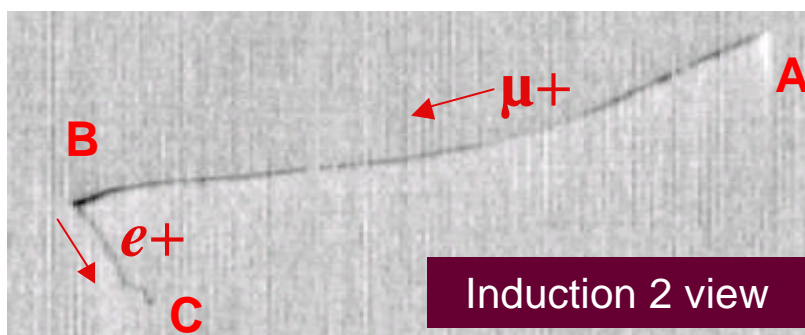
N.I.M. A516 (2004) 68.

**Study of the electron recombination in liquid Argon in the ICARUS TPC.**

To appear on N.I.M.

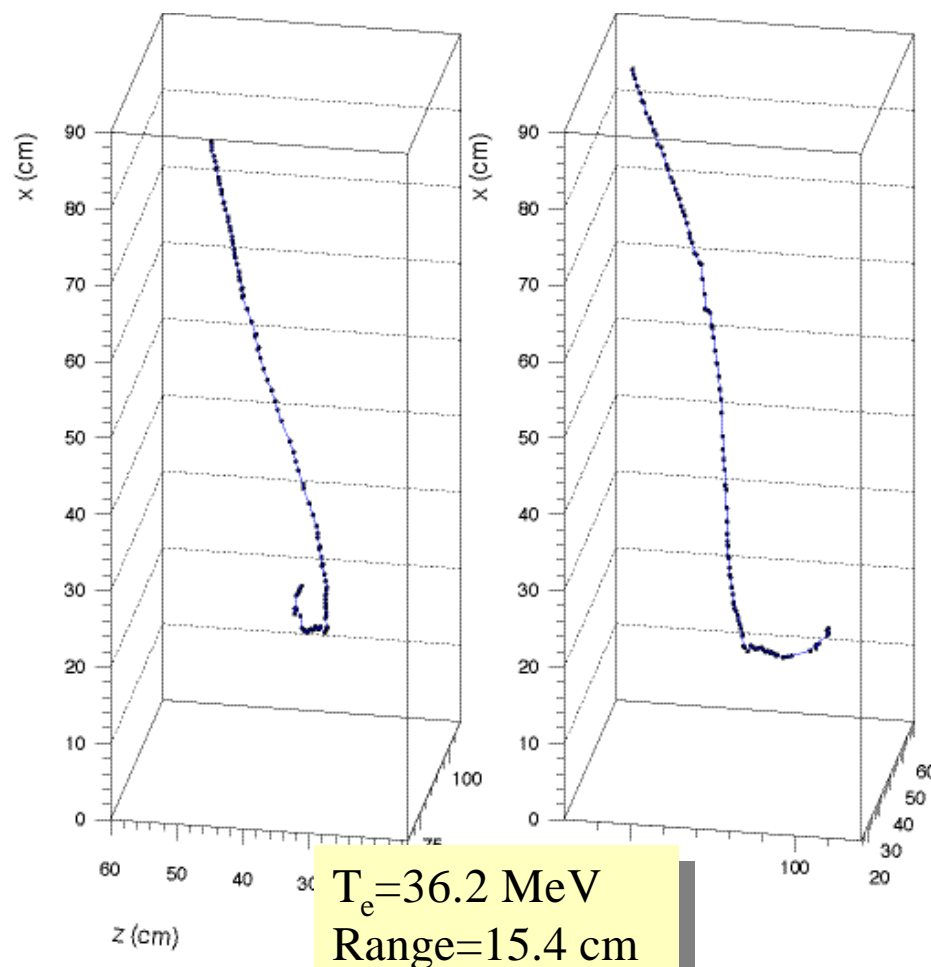
# 3D reconstruction stopping muon

$$m^+[AB] \rightarrow e^+[BC]$$



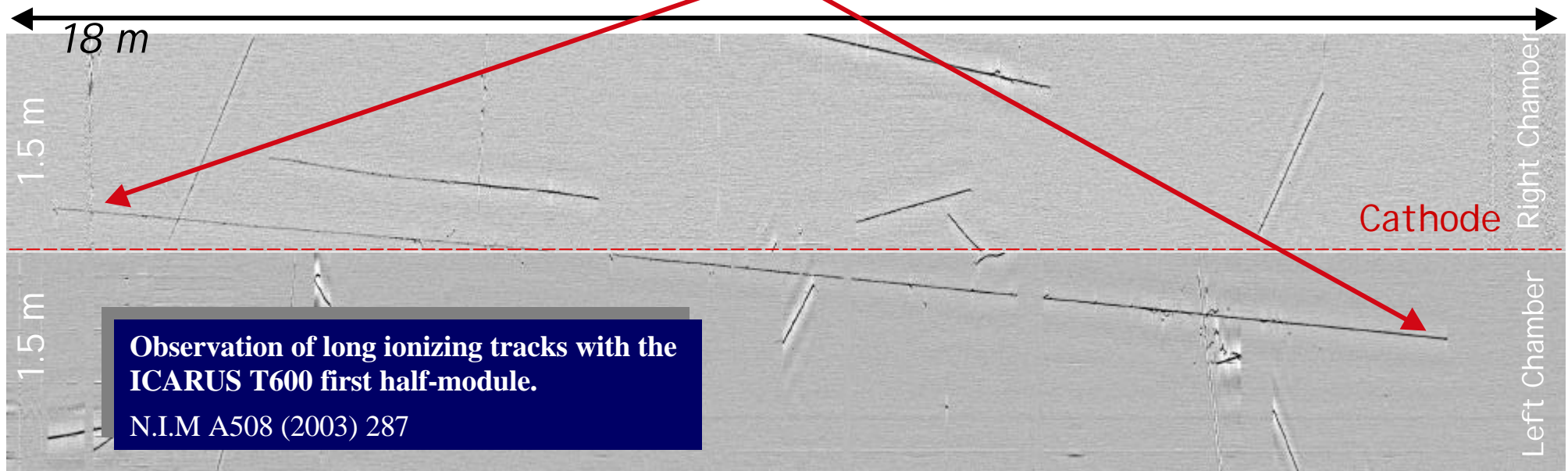
Measurement of the  $m$  decay spectrum with the ICARUS liquid Argon TPC.

Eur. Phys. Journ. C (2004). DOI 10.1140/epjc/s2004-01597-7

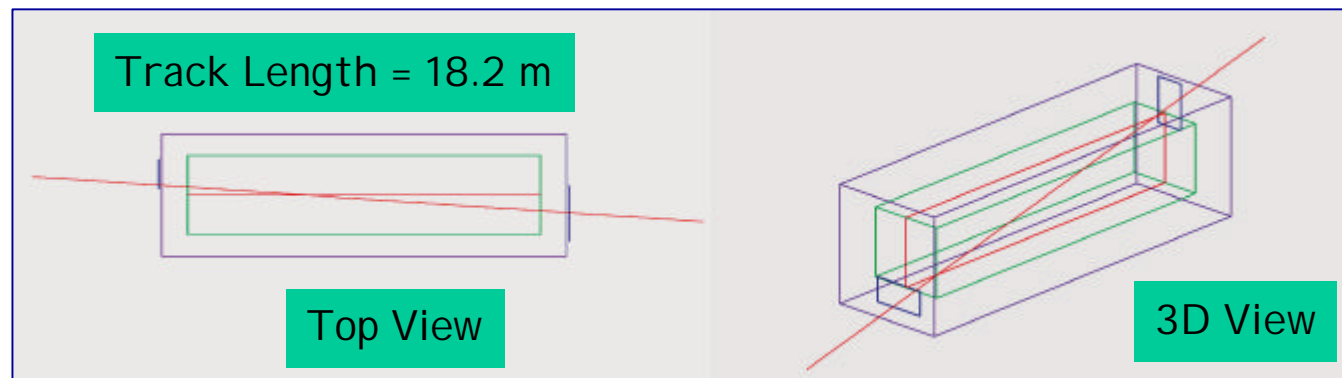




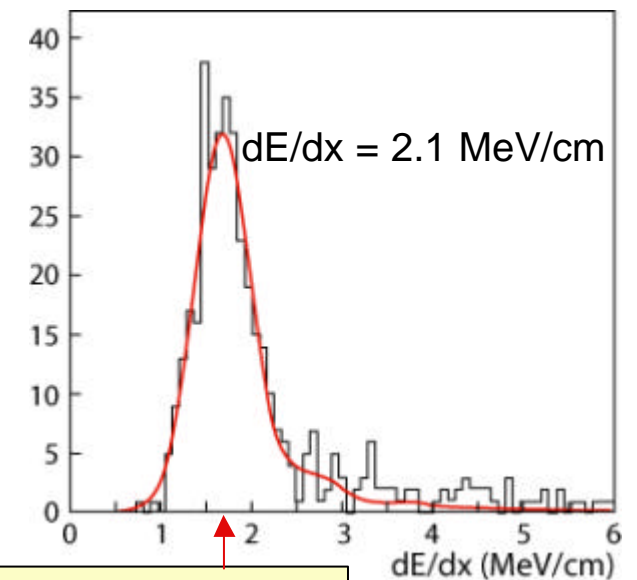
## Long longitudinal muon track crossing the cathode plane



~ 6000 points in each of the 3 readout planes !

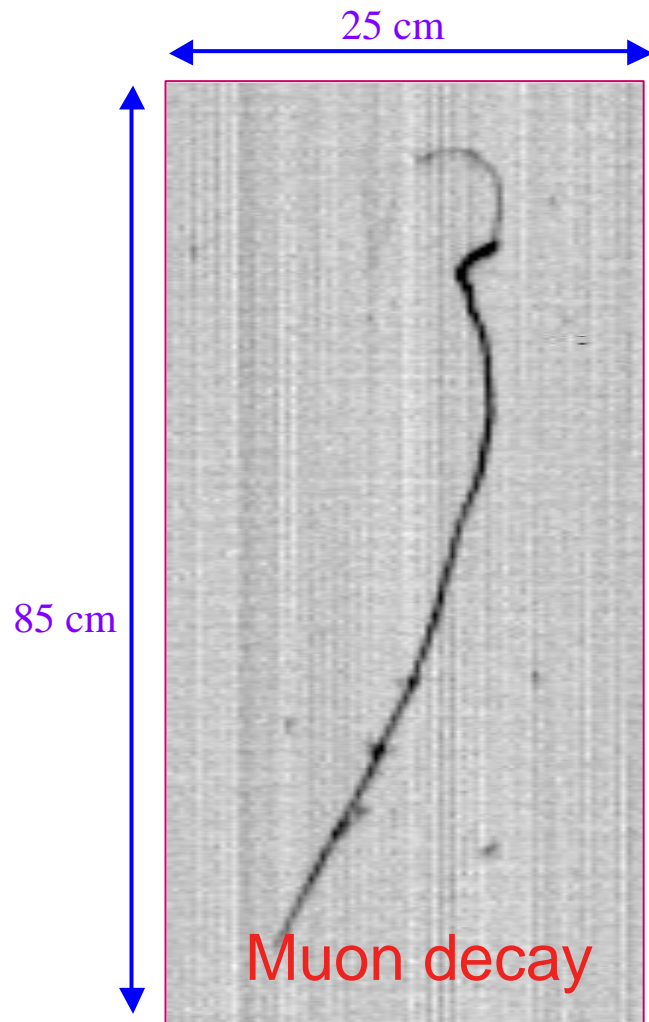


3-D reconstruction of the long track

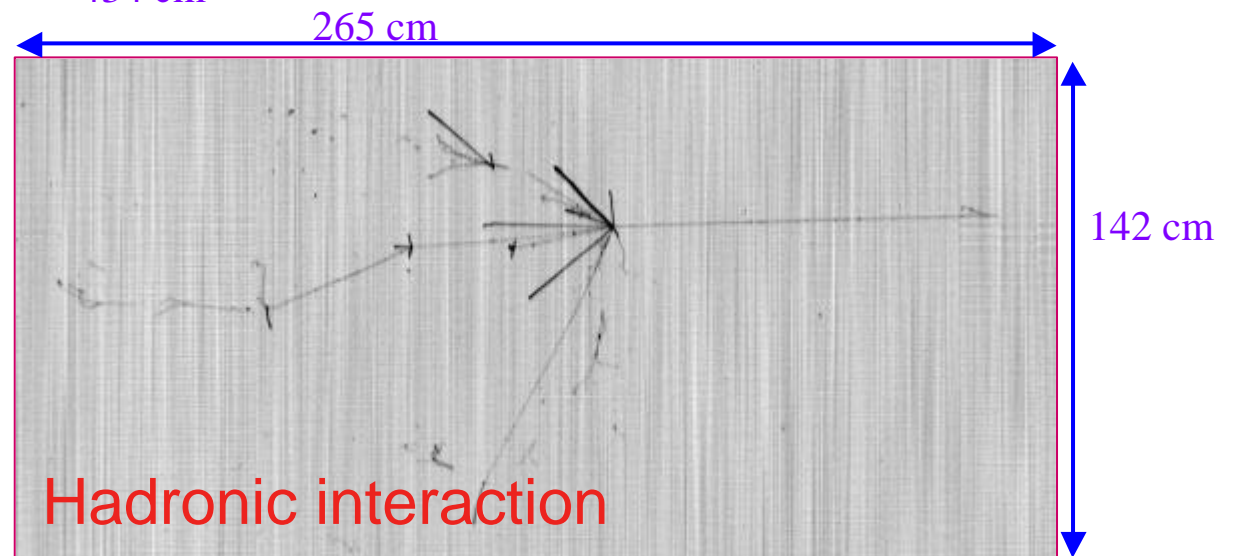
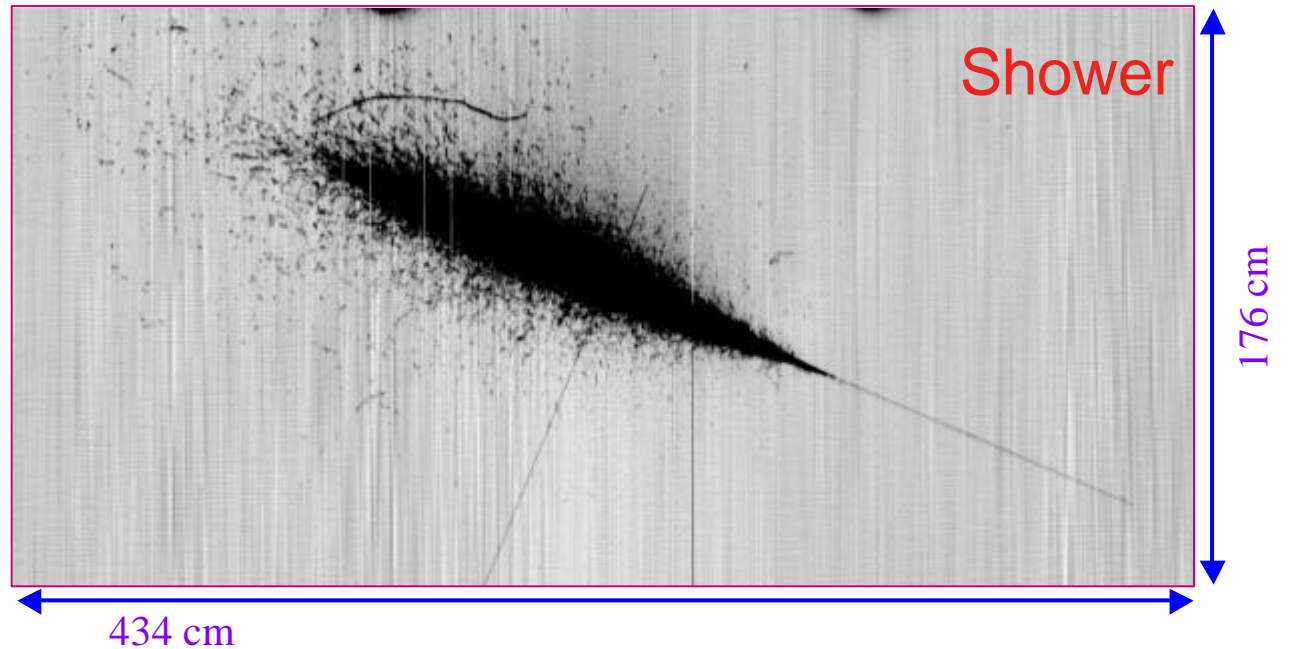


$dE/dx$  distribution along the track

# Cosmic ray interactions

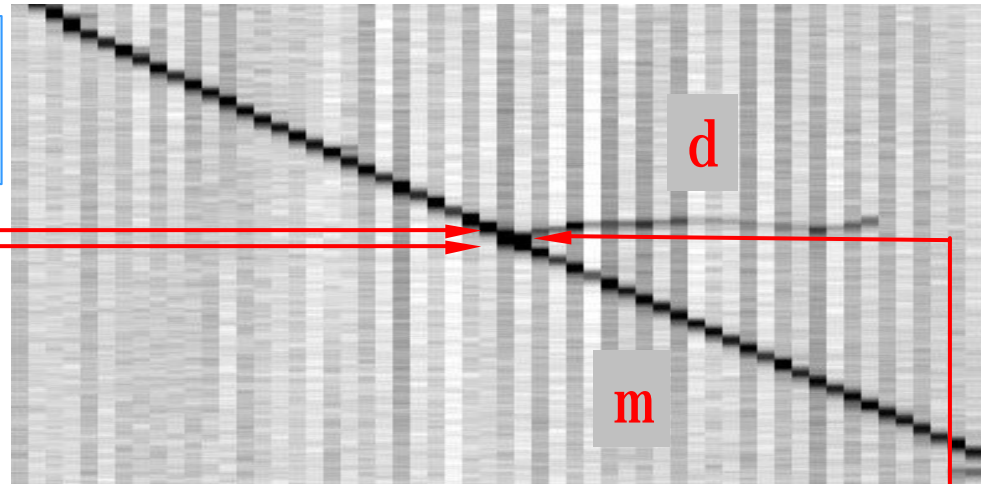


Run 960, Event 4 Collection Left



Run 308, Event 160 Collection Left

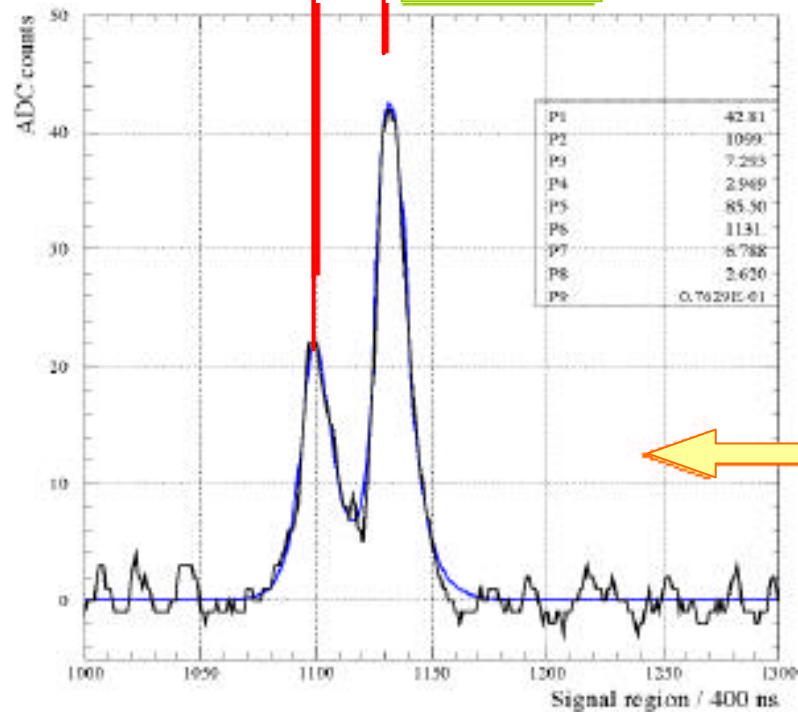
# Single wire performance



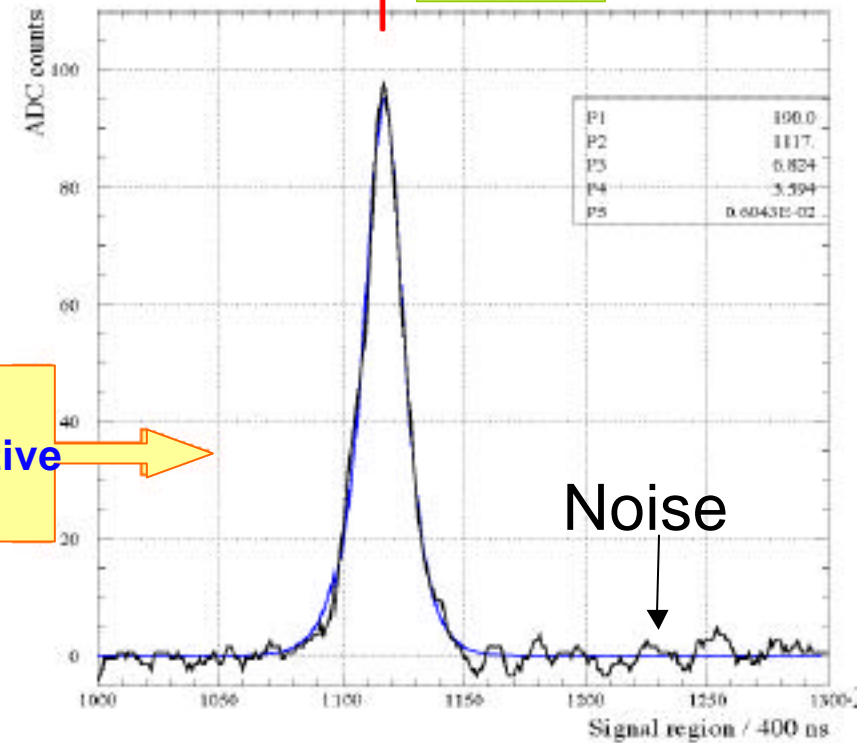
1.8 MeV

3.2 MeV

10 MeV



Two consecutive wires

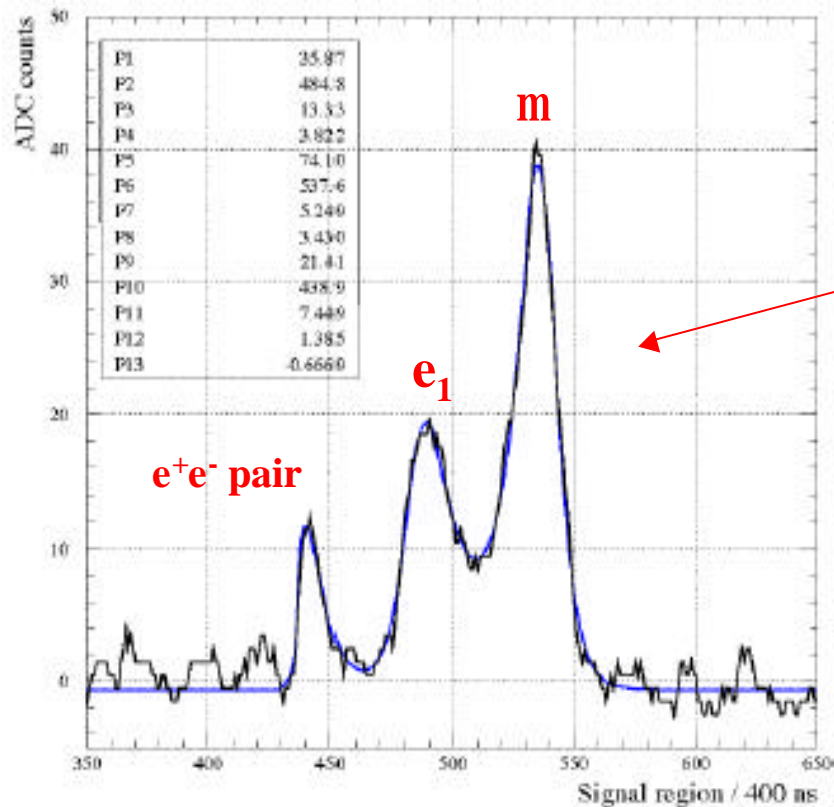


Noise

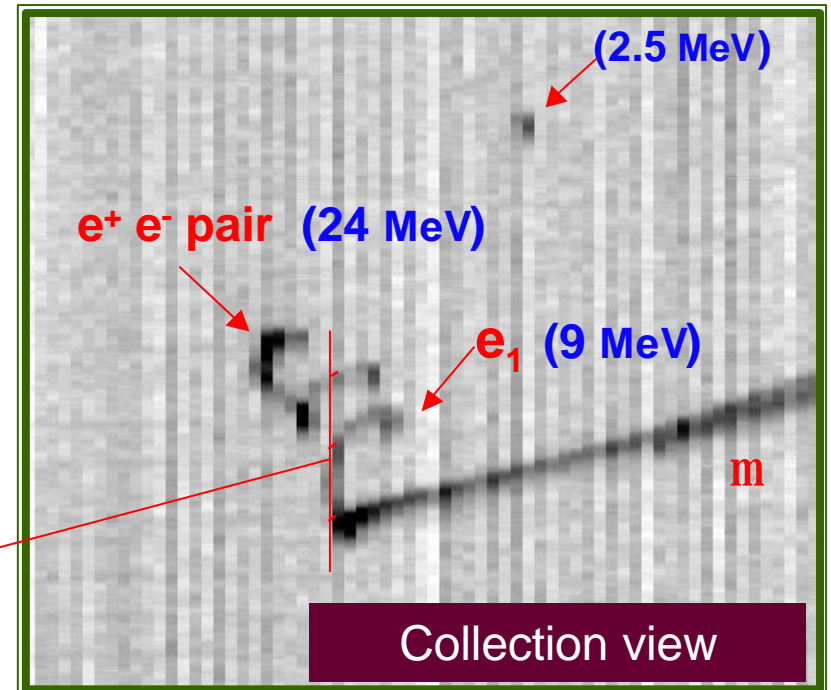


# Bremsstrahlung + pair-production

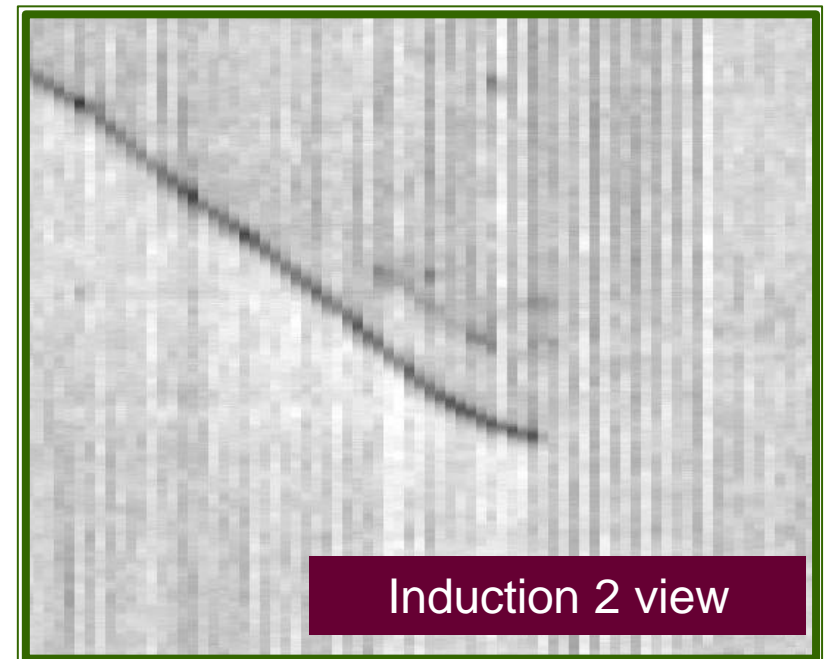
Run 975, Event 163



Fitted signal shapes  
on single wire



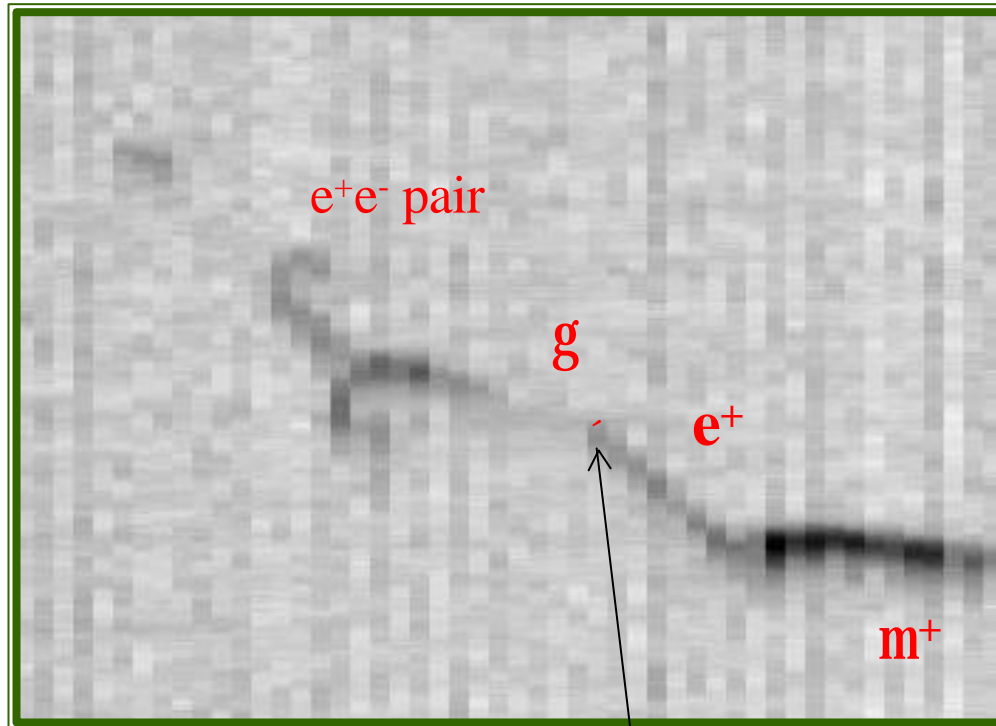
Collection view



Induction 2 view

# In-flight annihilation of positron

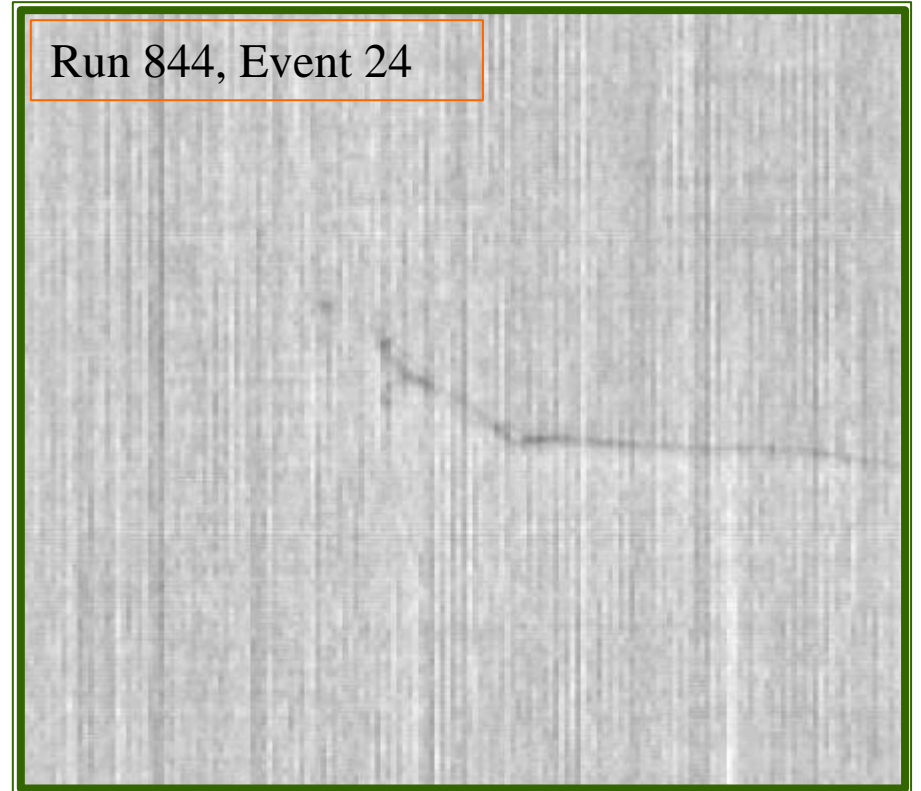
~ 20% of positrons from  $\mu$  decays expected to annihilate before stopping



Collection view

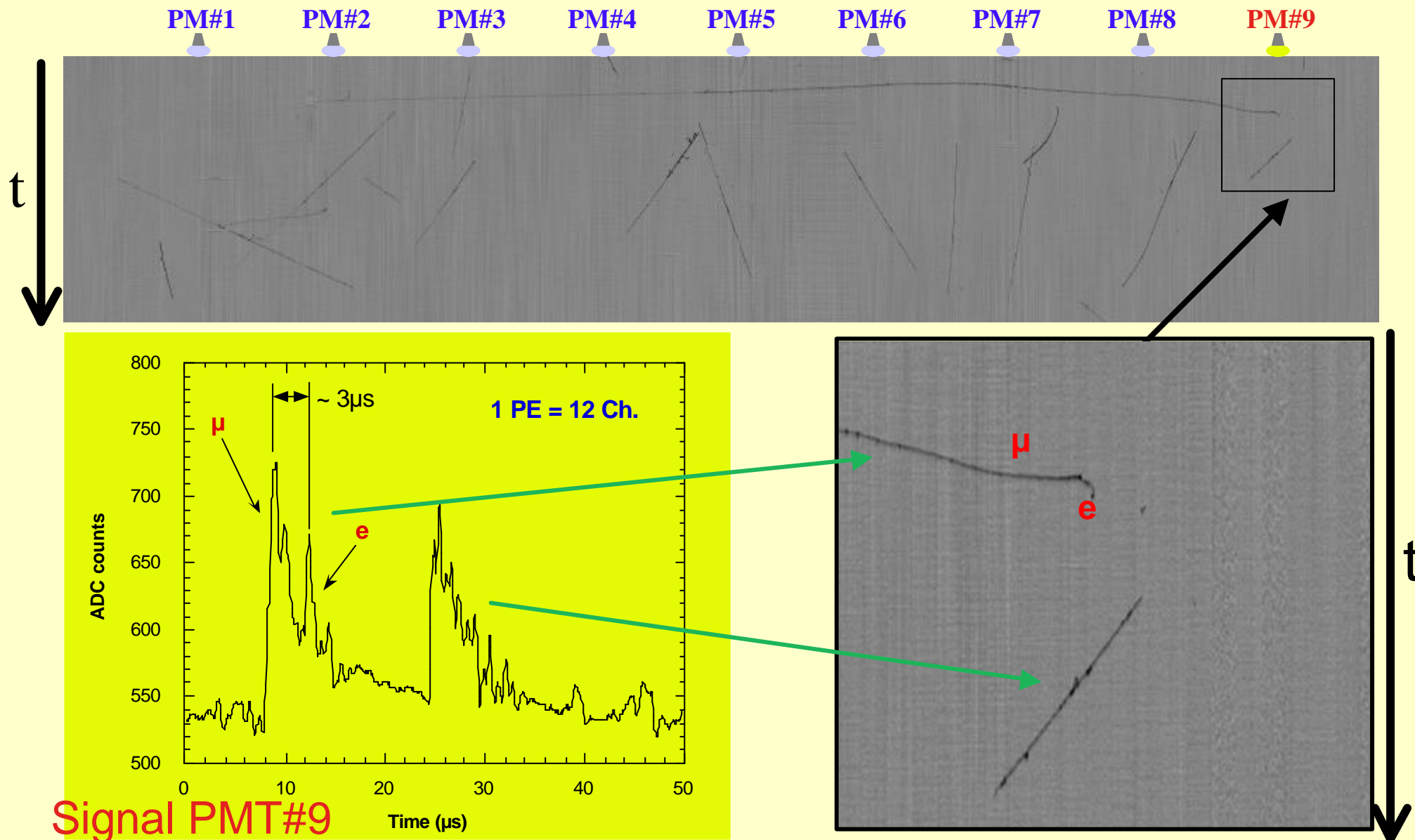
annihilation point

Run 844, Event 24



Induction 2 view

# Scintillation light readout





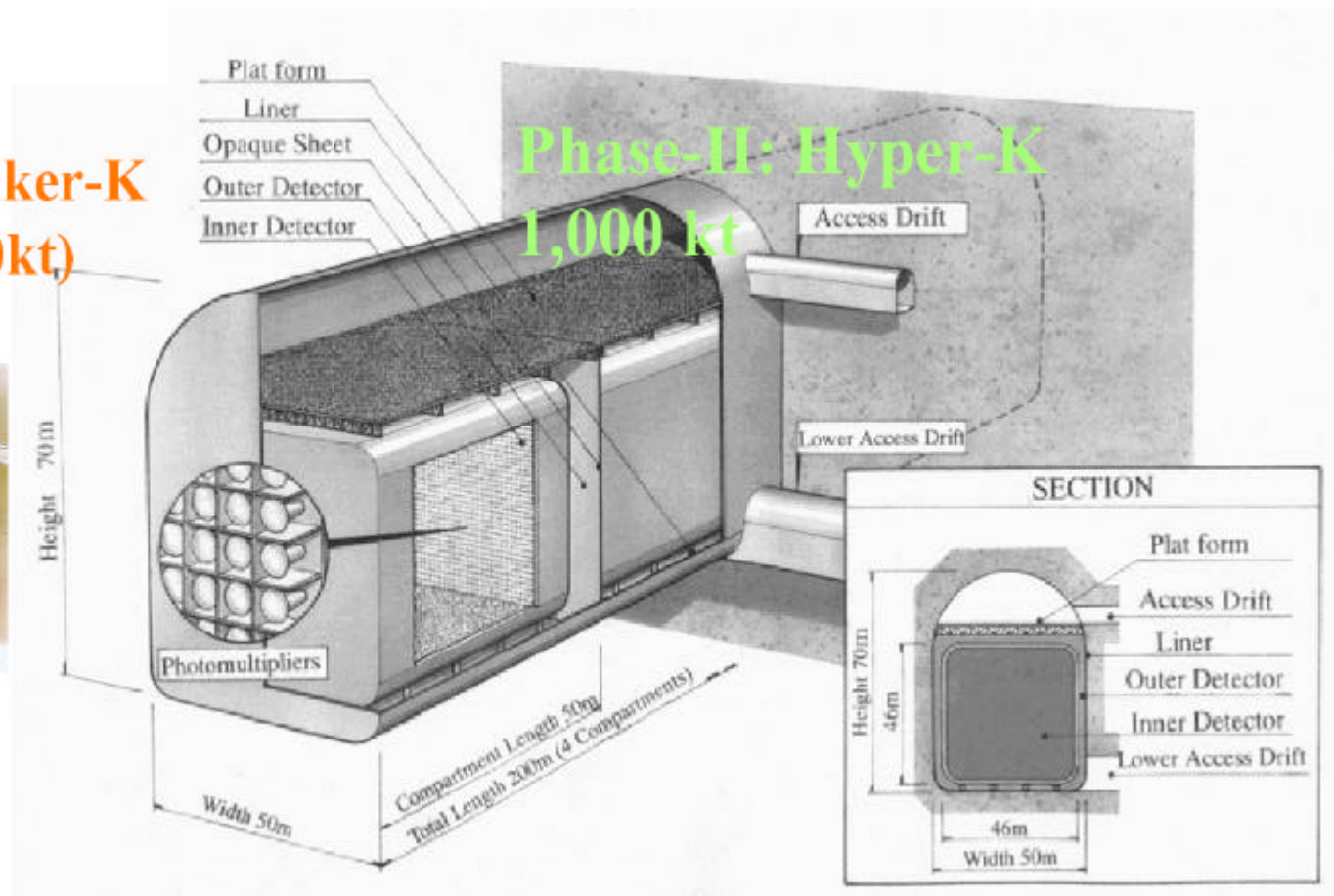


Towards the ultimate, very massive  
LAr detector for astroparticle  
physics experiments

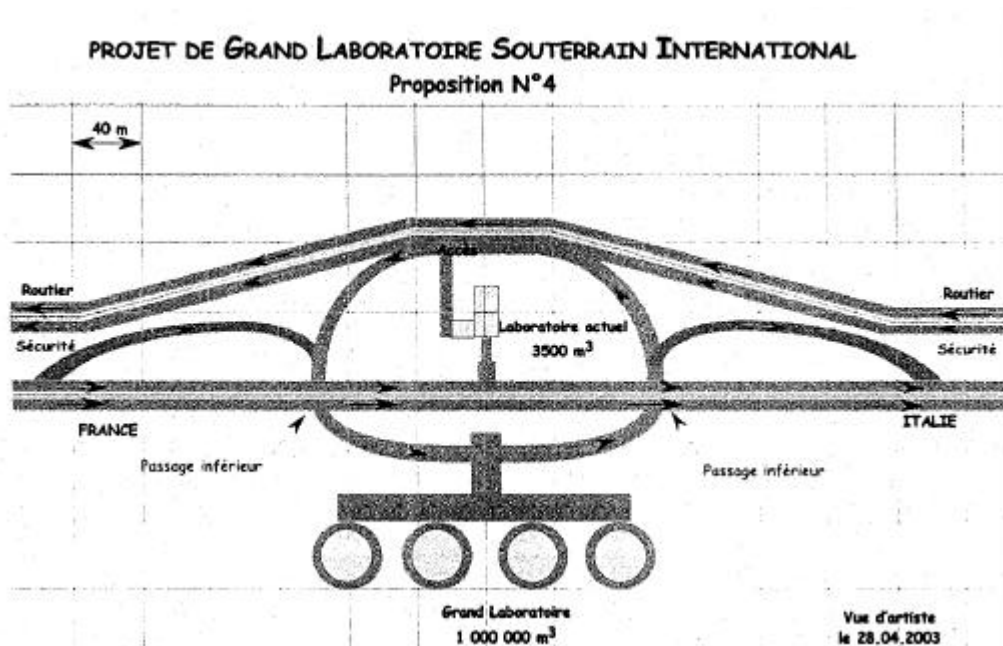
# Giant water Cerenkov (up to 1Mton)

- Widely perceived as a “straightforward” extension of SK (?)
- Many “proposals”, e.g., Hyper-K, UNO
- Many sites, e.g., Frejus, Kamioka, Homestake, etc.
- Physics case is “broad”: proton decay, neutrino properties, galactic supernovae, ...

**Phase-I: Super-K**  
**22.5kt (50kt)**



e.g.: the programme in Frejus: Superbeam or beta-beam from CERN (130 km)



**“Cooperation agreement”  
between  
French (IN2P3/CNRS, DSM/CEA)  
and Italian (INFN) Institutions**

.....

« The DSM, IN2P3 and the INFN agree to prepare the design of a very Large Underground Laboratory in the new Fréjus tunnel, with complementary features with respect to the Gran Sasso laboratory, to be submitted as a joint proposal to the French and Italian governments.

The institutions aim at associating the Fréjus and Gran Sasso laboratories in a single entity, a European Joint Laboratory, open to the world scientific community to carry out advanced experiments in particle, astroparticle and nuclear physics in the coming decades, on topics such as matter stability, neutrino mixing and mass, stellar collapses and nuclear astrophysics »

Detector options: 4 x 250 kton Cerenkov tanks ? ... a 100-200 kton LAr detector ?

**OTHER SITES MAY WELL BE CONSIDERED FOR VERY LARGE DETECTORS**

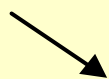
- Astroparticle physics (without beam) interesting 'per se'
- Investment for the detector may be comparable to the beam facility
- Existing underground sites like mines, etc.



# An effective alternative: a very massive LAr detector

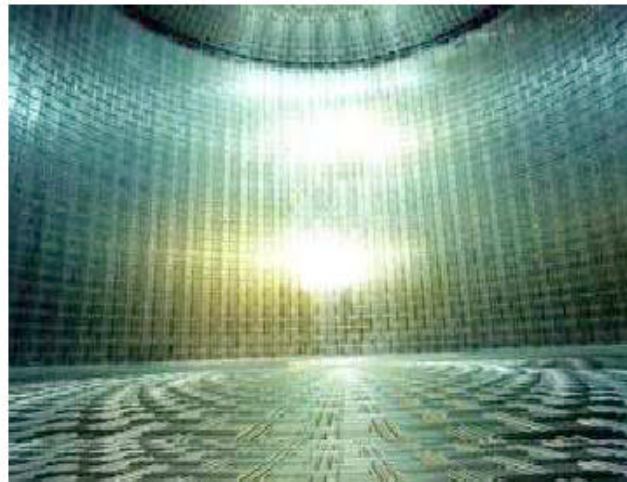
- Can one extrapolate LAr technology to  $> 100$  kton detector?
- Ultra-pure liquid Argon is cheap and industrially available.
- Cryogenic tanks for Liquefied Natural Gas (LNG) are industrially well developed products with considerable sizes. Heat leaks are small ( $\sim 5$  W/m<sup>2</sup>)
- Ideas being developed:

See e.g.:



**Experiments for CP violation: a giant liquid Argon scintillation, Cerenkov and charge imaging experiment.**

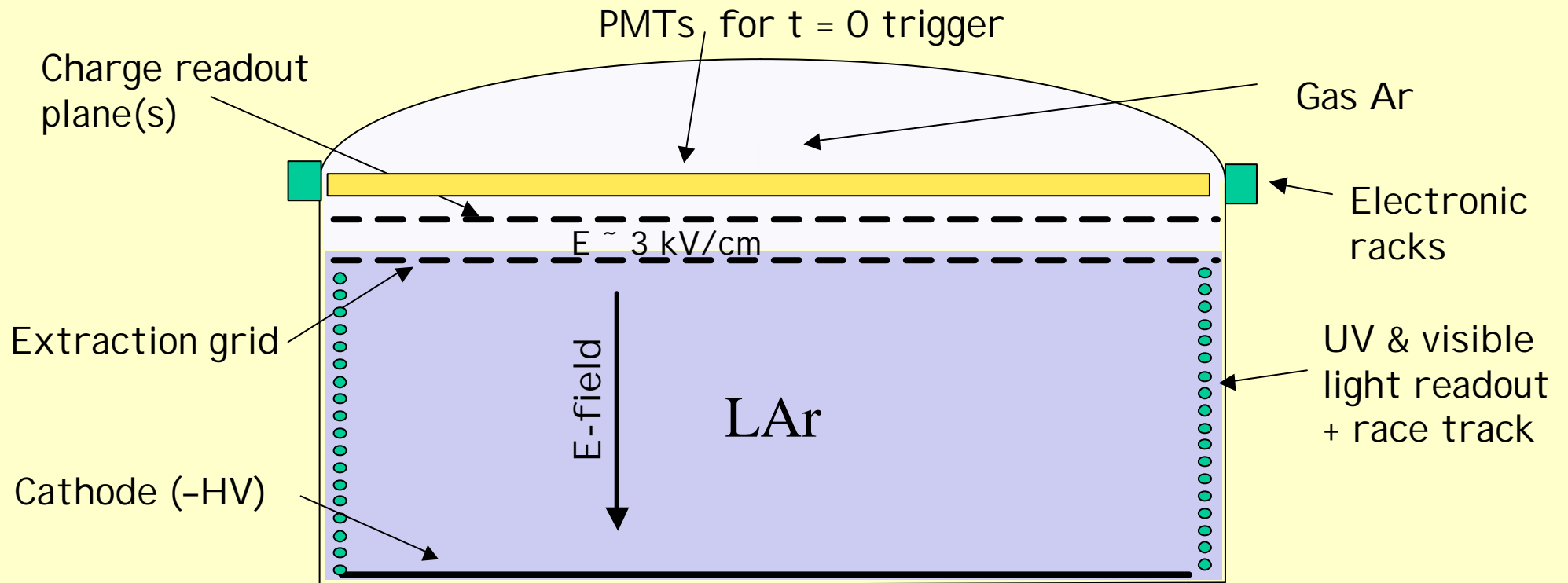
A.Rubbia, Proc. II Int. Workshop on Neutrinos in Venice, 2003, hep-ph/0402110



**Liquid Argon Imaging Technology.**

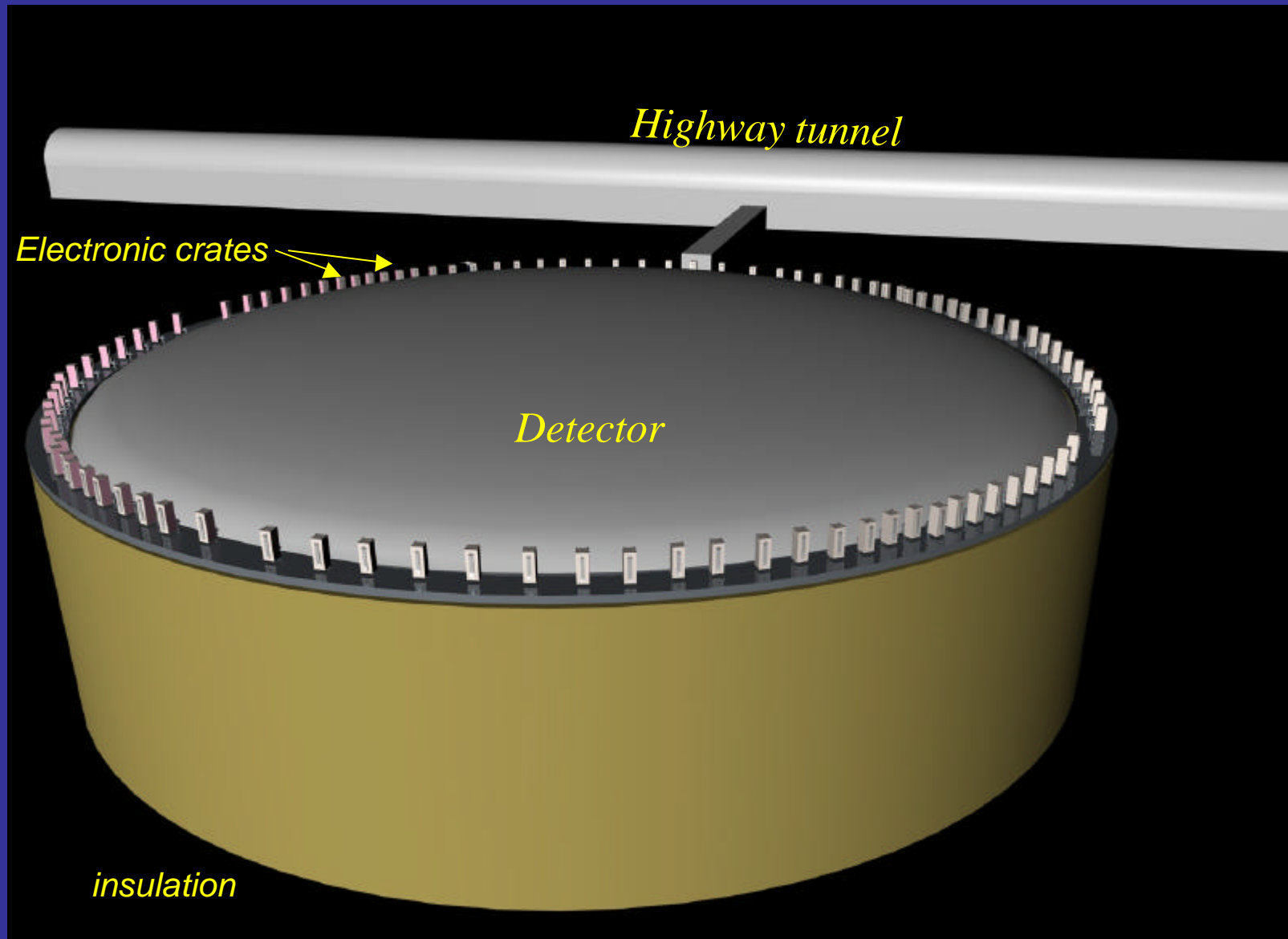
C.Rubbia, talk at the SLAC Experimental Seminar February 17, 2004.

# A 100-200 kton LAr detector: conceptual design

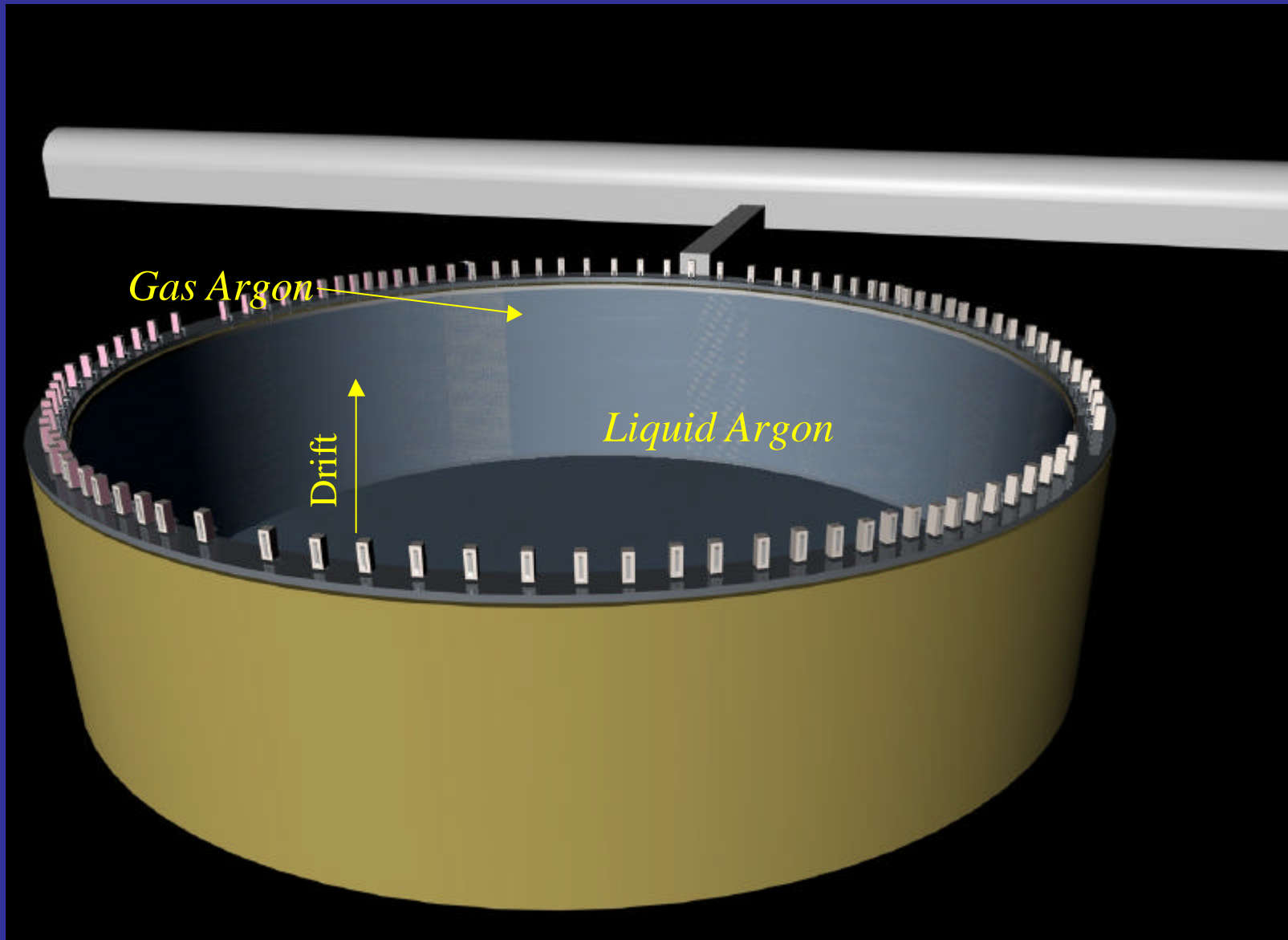


- 70-120 m diameter, 15-20 m height,  $\sim 100000$  readout wires
- In situ cryogenic plant,  $5 \text{ W/m}^2$  heat input, continuous re-circulation (purity)
- Filling speed (100 kton): 150 ton/day  $\rightarrow$  2 years to fill
- $E = 500\text{-}1000 \text{ V/cm}$   $V_0 = 750\text{-}1500 \text{ kV}$ , drift time  $\sim 10 \text{ ms}$
- Max charge attenuation/required multiplication:  $\sim$  factor 100 (e.g. thin wires/pads)
- Same detector: charge imaging, scintillation, Cerenkov light

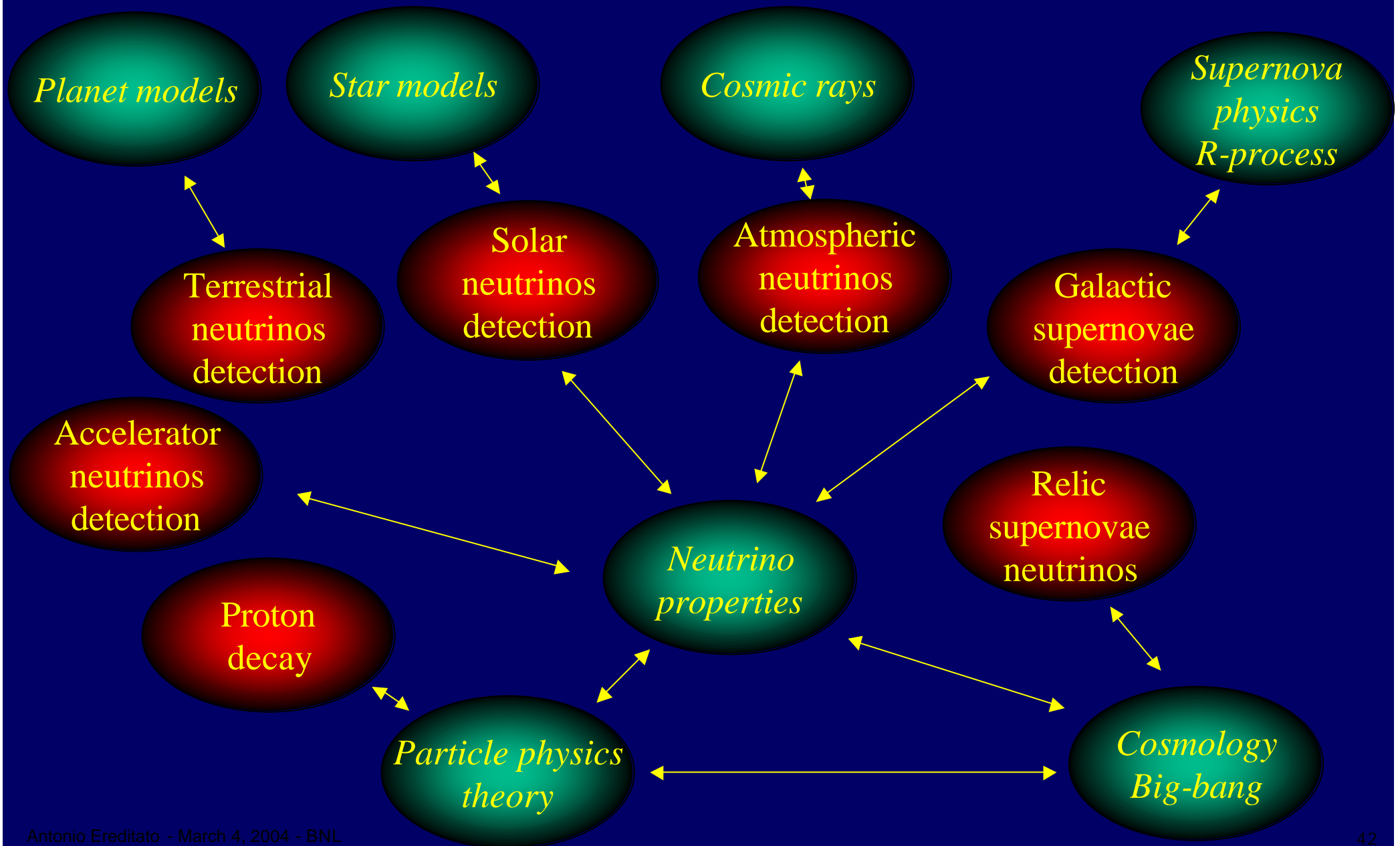
## Highway tunnel laboratory option



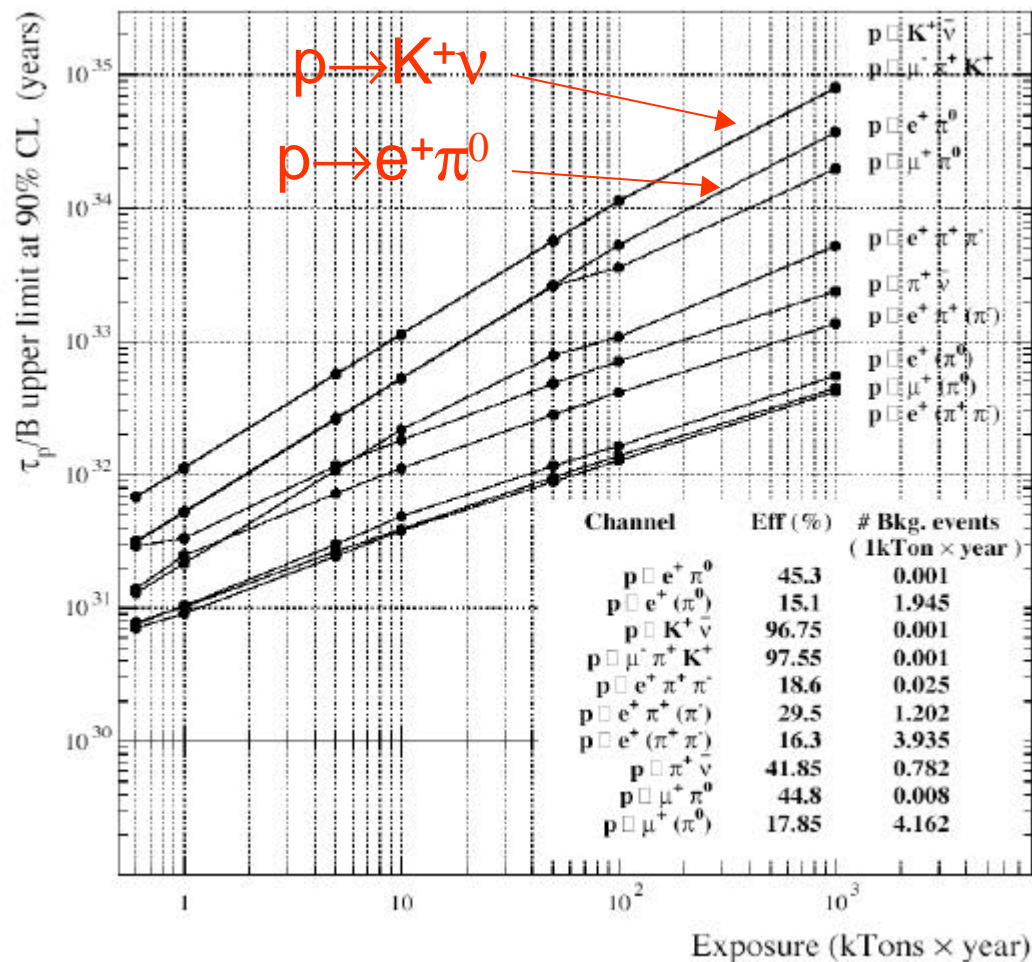




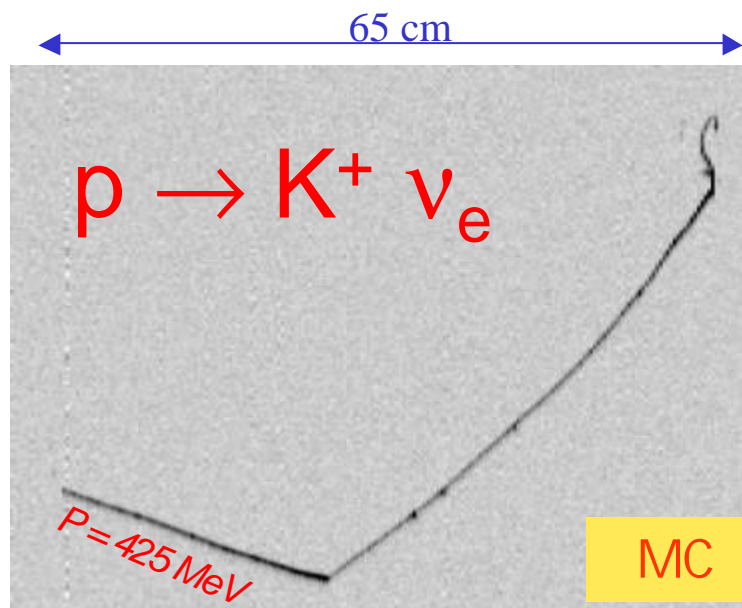
# Physics



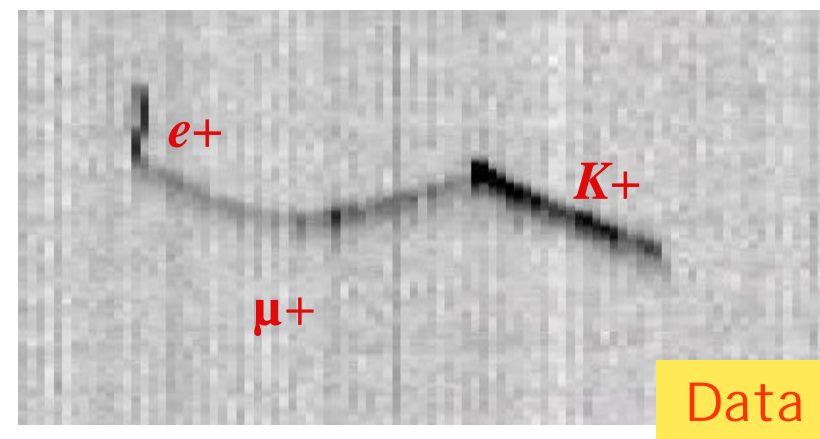
# Proton decay: sensitivity vs exposure



$6 \times 10^{34}$  nucleons (100 kton)  $\Rightarrow$   
 $t_p / Br > 2 \times 10^{34} \text{ years} \times T(\text{yr})^{-1} \text{ @ 90 CL}$



*"Single" event detection capability*



T600: Run 939 Event 46

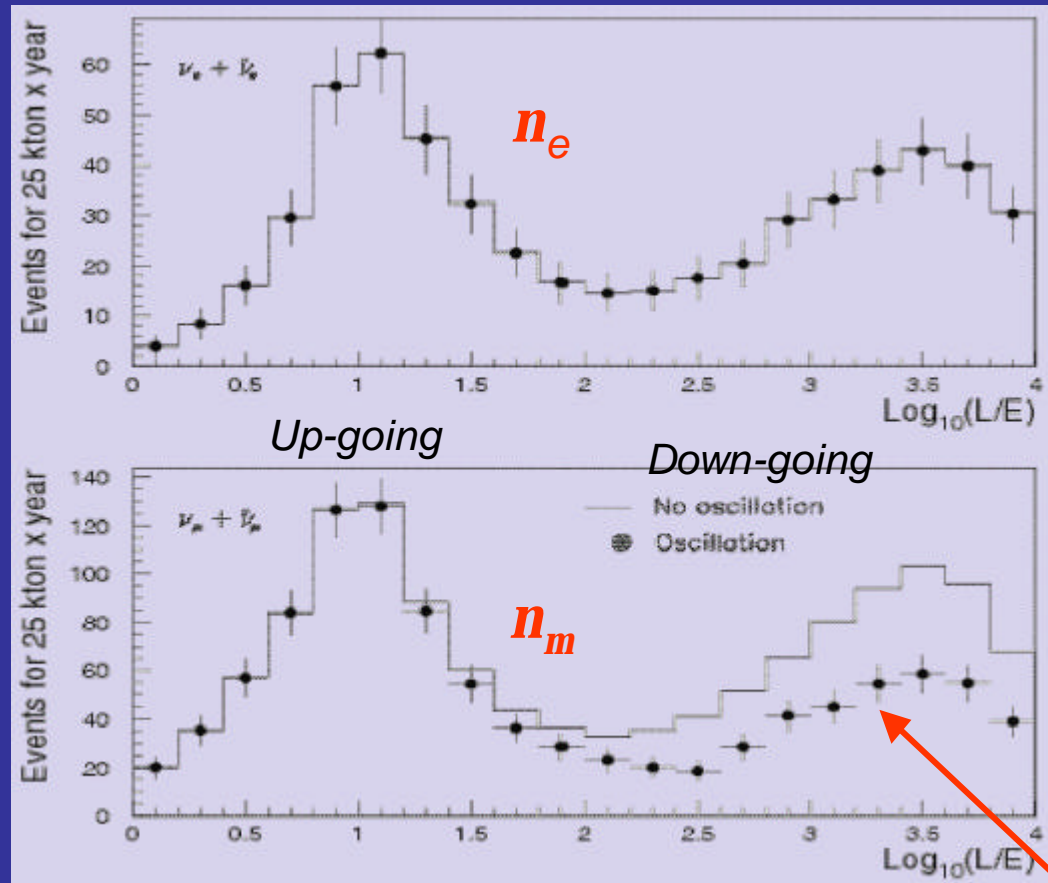


# Atmospheric neutrino events

The LAr detector will provide:

- ✓ Observation of atmospheric neutrino events with high quality
- ✓ Unbiased, systematic free observation
- ✓ Excellent energy and angular reconstruction

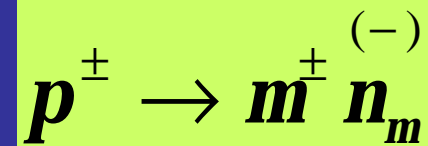
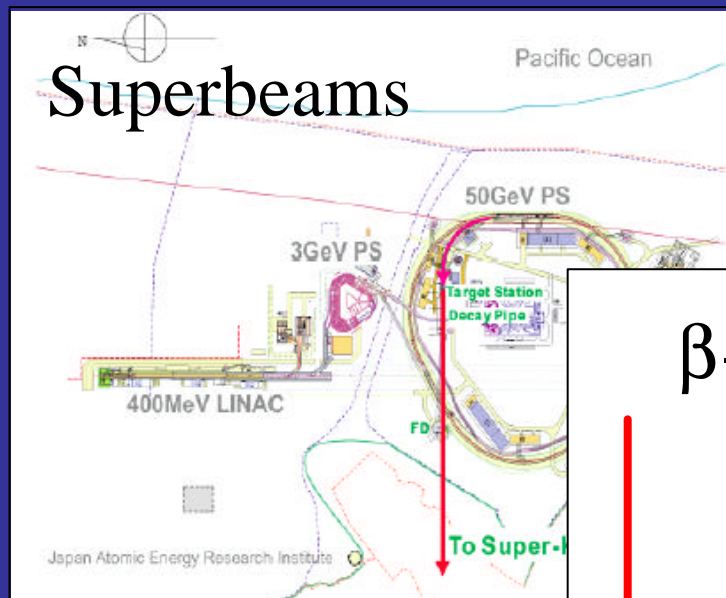
Event rate in 1/10 of a year for 200 kton  
8 years of ICARUS T3000 operation



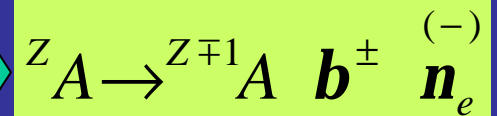
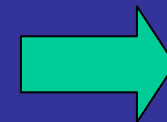
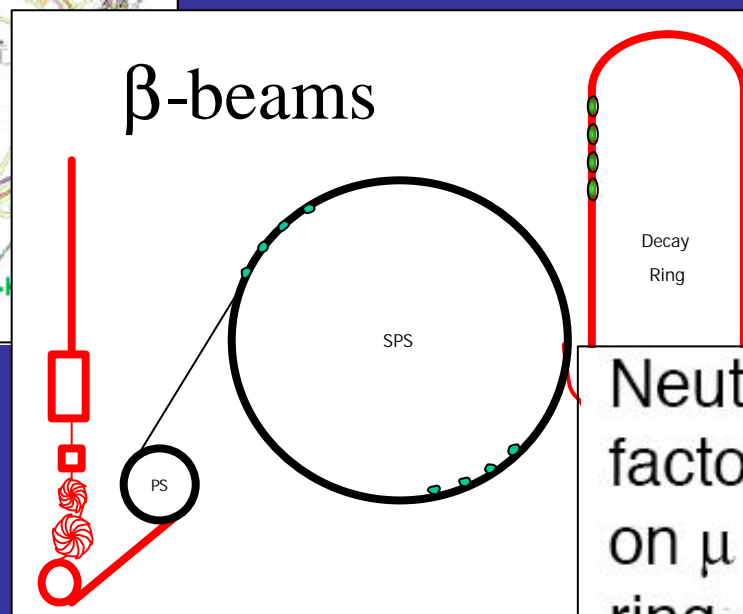
$\log_{10}(\text{length/energy}) \rightarrow$

Effect of oscillations

# Accelerator neutrinos

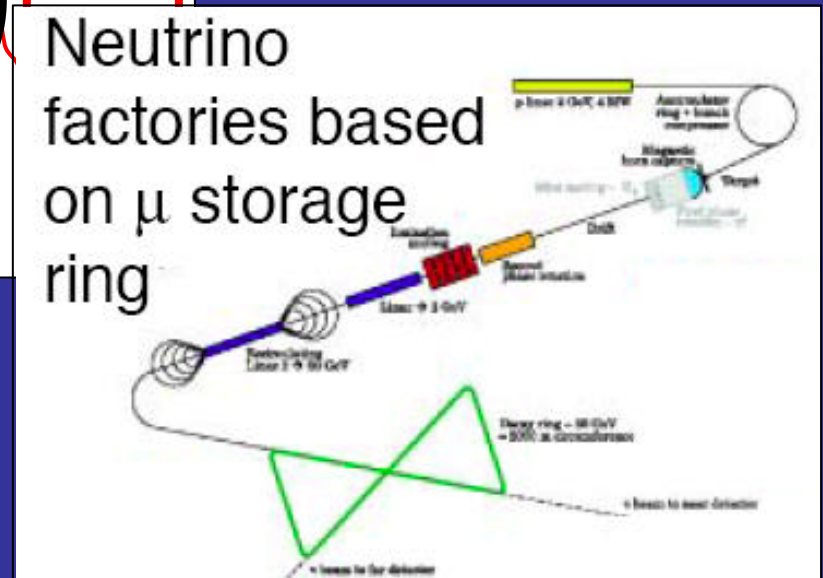
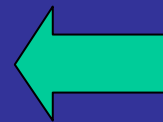
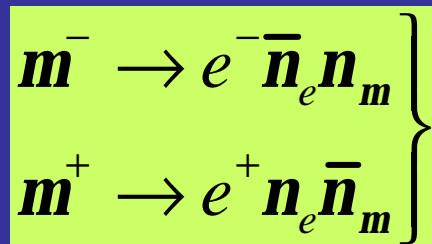


*Select focusing sign*



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## Events for 100 kton detector mass

Number of targets for nucleon stability:

$$6 \times 10^{34} \text{ nucleons} \Rightarrow \tau_p/\text{Br} > 10^{34} \text{ years} \times T(\text{yr}) \times \varepsilon @ 90 \text{ CL}$$

Low energy Super-Beams or beta-beams:

$$460 \nu_\mu \text{ CC per } 10^{21} \text{ 2.2 GeV protons (real focus) @ } L = 130 \text{ km}$$

$$15000 \nu_e \text{ CC per } 10^{19} \text{ }^{18}\text{Ne decays with } \gamma=75$$

Atmospheric:

$$10000 \text{ atmospheric events/year}$$

$$100 \nu_\tau \text{ CC /year from oscillations}$$

Solar:

$$324000 \text{ solar neutrinos/year @ } E_e > 5 \text{ MeV}$$

Supernova type-II:

$$20000 \text{ events @ } D=10 \text{ kpc}$$



## Conclusions

- Ultra-pure liquid Argon is now well mastered industrially. Events of remarkable quality, excellent energy and 3D angular reconstruction are routinely produced. The detector is self-triggering and continuously sensitive.
- A 600 ton (ICARUS T600) detector has been operational since a couple of years and it is going to be commissioned at LNGS for underground operation. R&D completed and industrialization phase addressed. Real events already available.
- ICARUS at LNGS: the most important milestone for the technology; test-bed with 3 kton of liquid Argon in a difficult underground environment: simultaneous detection of cosmic events and CNGS beam neutrinos.
- Signal multiplication after extraction in the gas and the industrial technology of LNG allow to conceive a very massive ( $\sim 100$ -200 kton) LAr detector for cosmic and LBL neutrinos, at a cost comparable to a Mton water Cerenkov detector but with an unparalleled spatial and energy resolutions and event reconstruction capabilities.